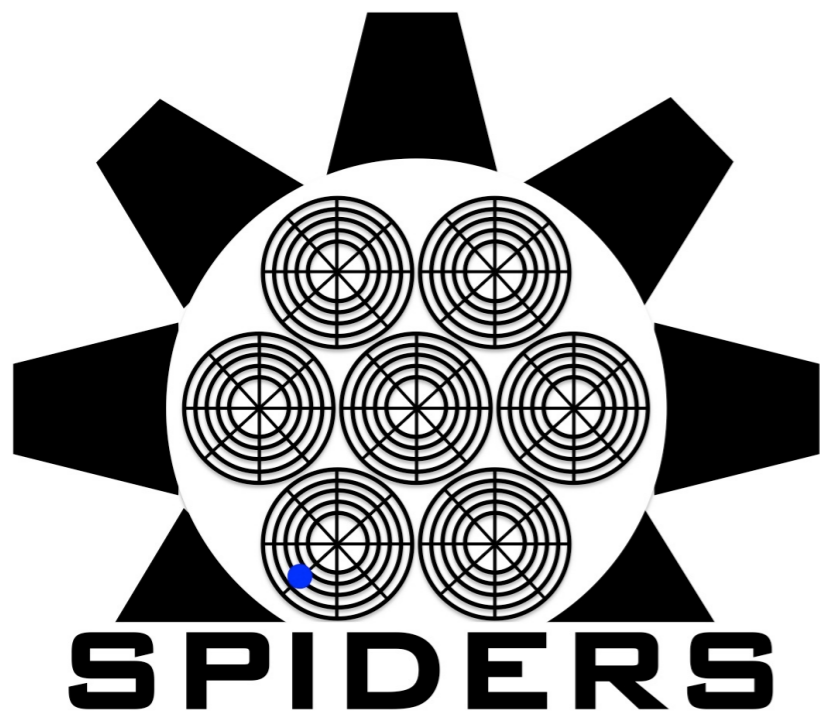
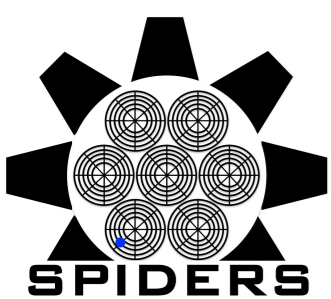


eBOSS/SPIDERS:AGN Target selection and first results



Tom Dwelly (MPE)
for the SPIDERS team



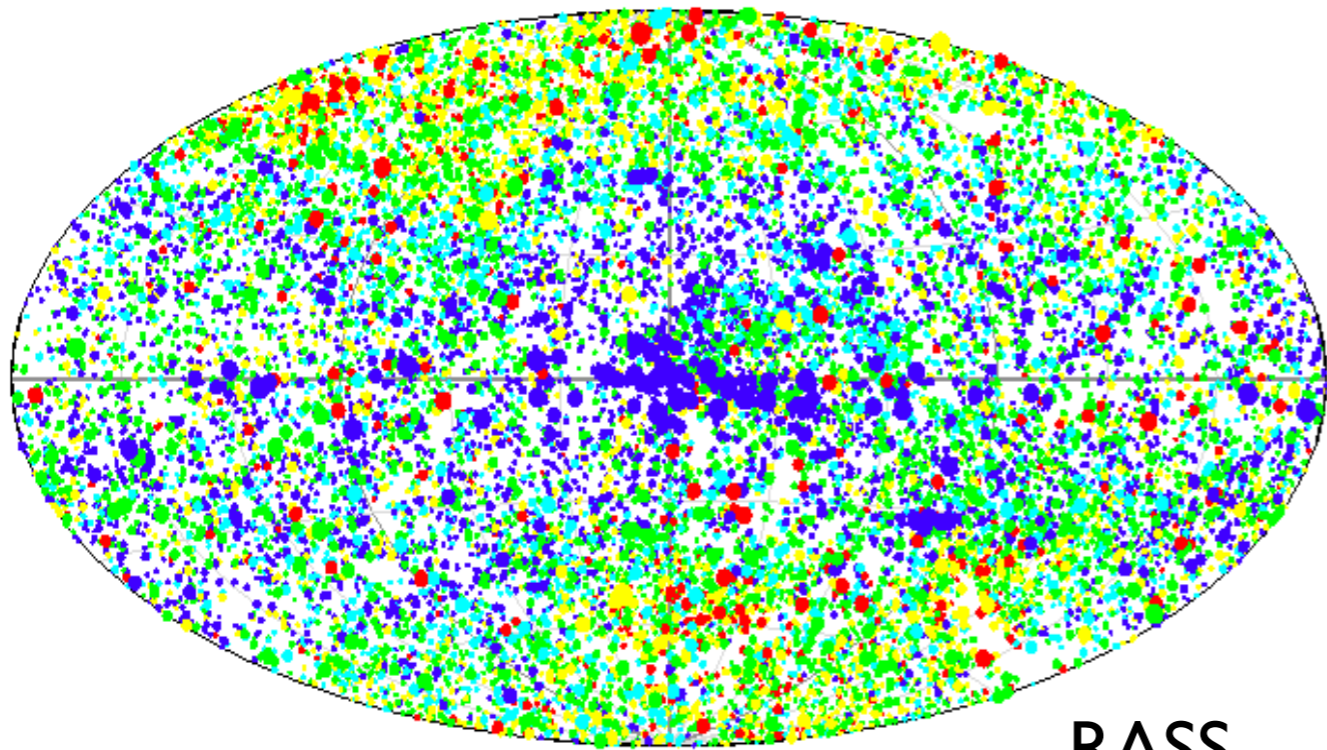
Outline



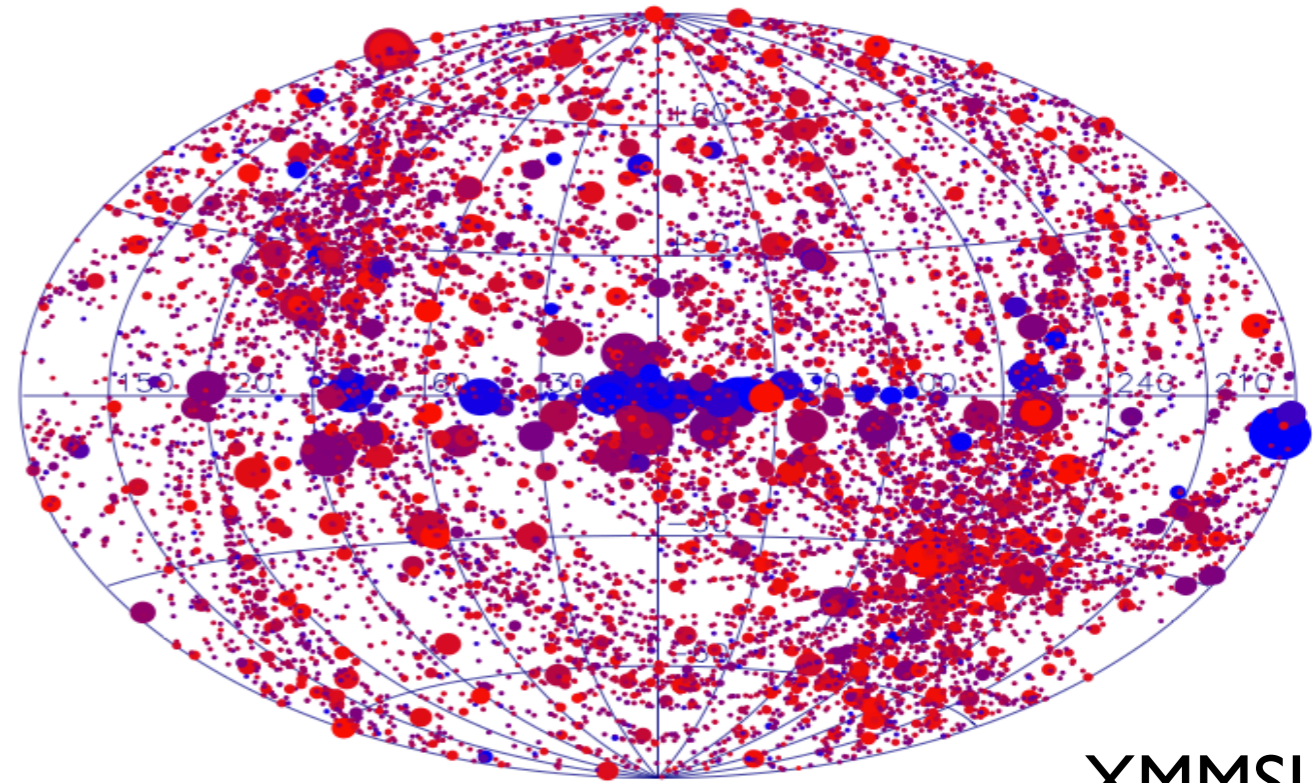
SPIDERS: Spectroscopic Identification of (pre-)eROSITA Sources

X-ray selected AGN component

- Pre-eROSITA targets (RASS, XMMSL)
- Counterpart selection method using WISE
- Sky coverage and target numbers
- First results from eBOSS
- SEQUELS results (if enough time)



RASS



XMMSL

SPIDERS AGN in 'Tier-0'



Targets Selected from
RASS and XMMSL
Available at start of eBOSS
operations (July 2014)



- Parent footprint:

BOSS imaging = 10800 deg^2

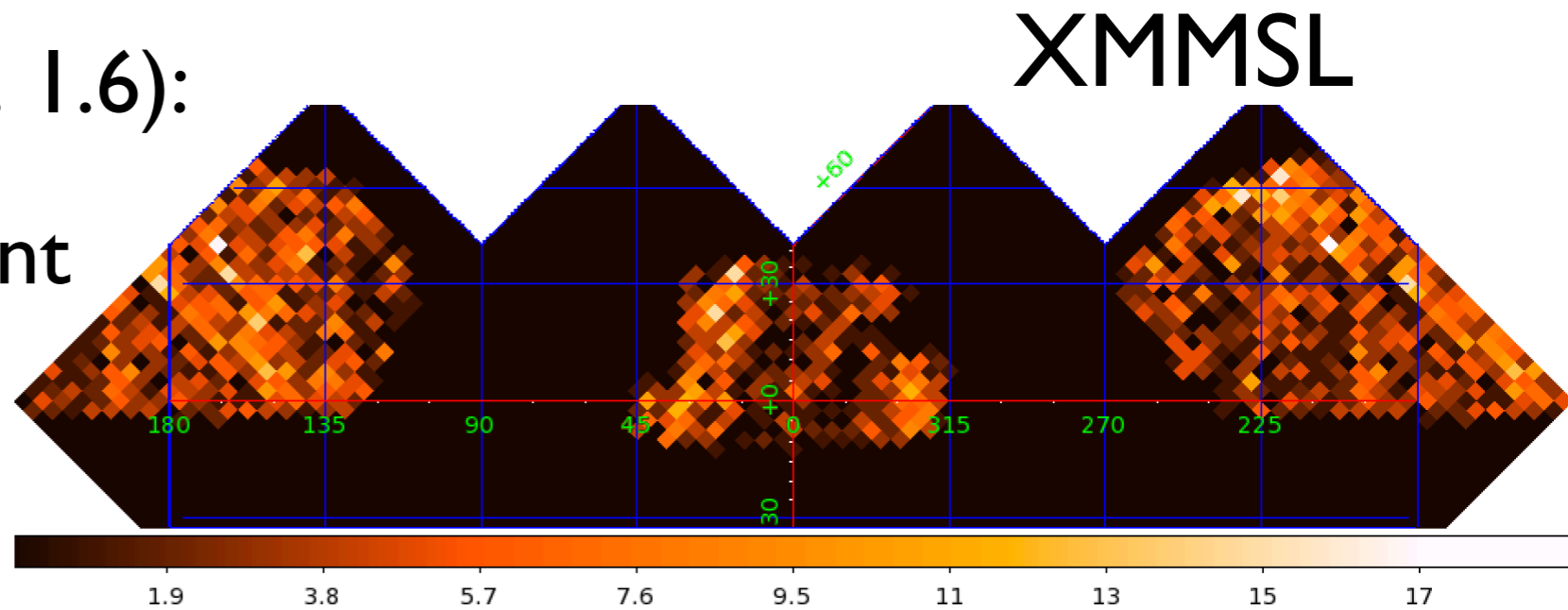
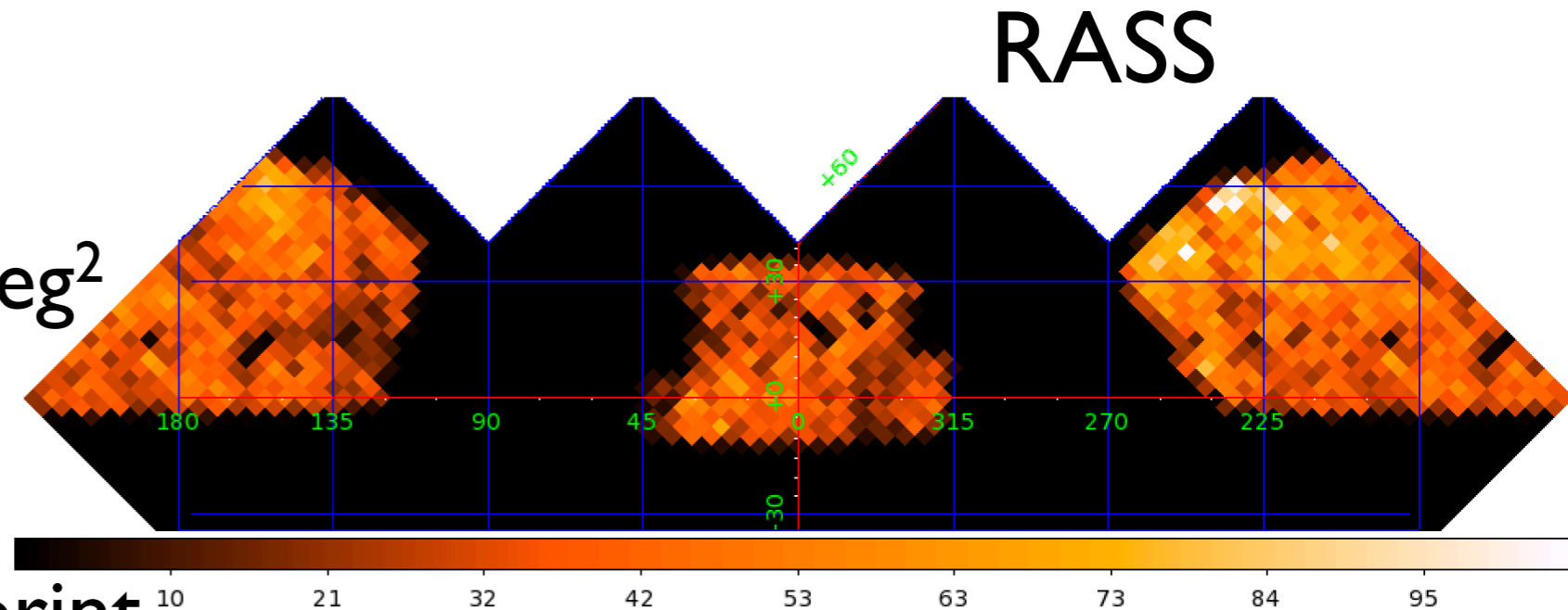
- RASS (BSC+FSC):

32400 sources in footprint

- XMM Slew Survey (rel. 1.6):

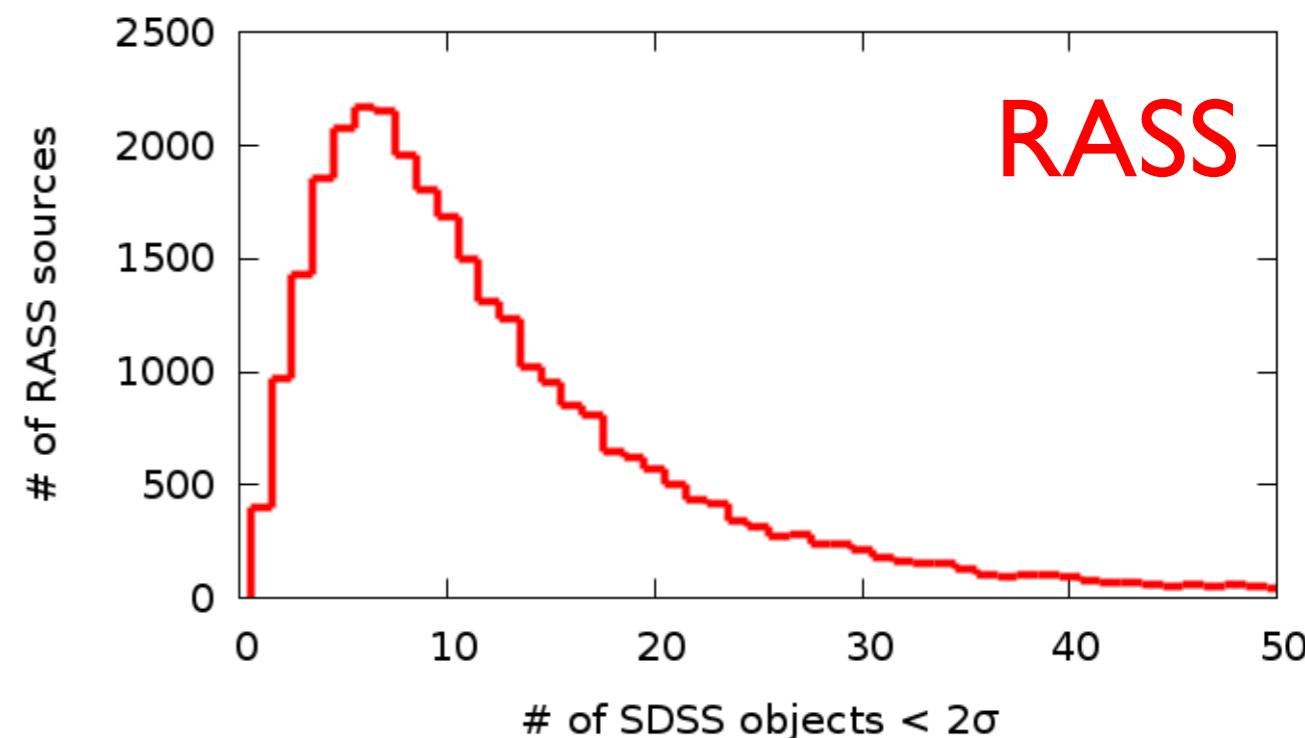
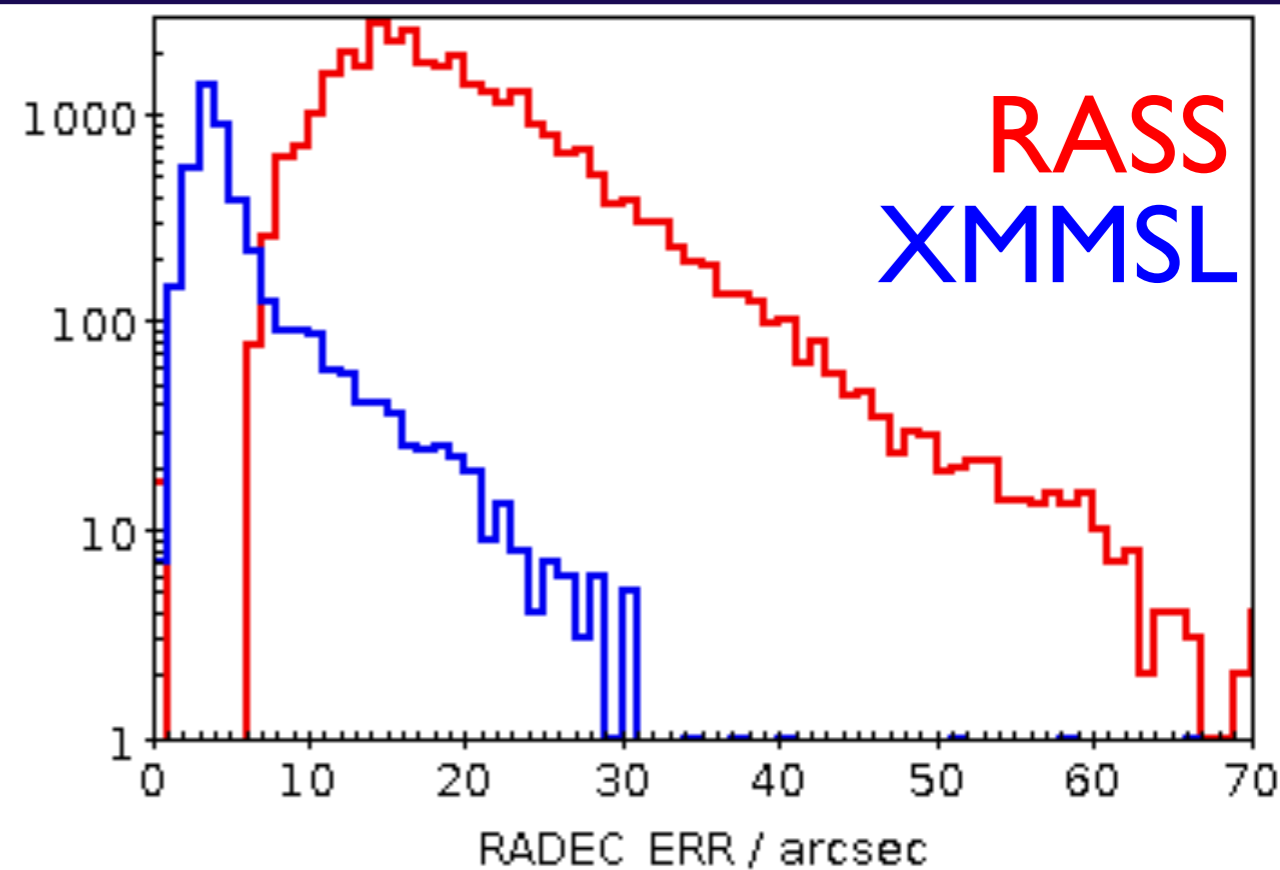
4300 sources in footprint

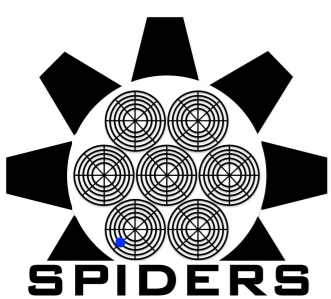
- Some sources appear in both RASS and XMMSL



Maps: HEALPIX $n_{\text{side}}=16$ ($13.4 \text{ deg}^2 \text{ pixel}^{-1}$)

- X-ray position errors:
 RASS: $\langle 1\sigma \rangle = 20''$ (95%ile=35'')
 XMMSL: $\langle 1\sigma \rangle = 6''$ (95%ile=15'')
- Heterogeneous mix of AGN, Stars, Clusters etc
- ~ 8 objects arcmin⁻² at SDSS imaging depths ($g' \sim r' \sim 22$ AB)
- How can we choose the correct counterpart?





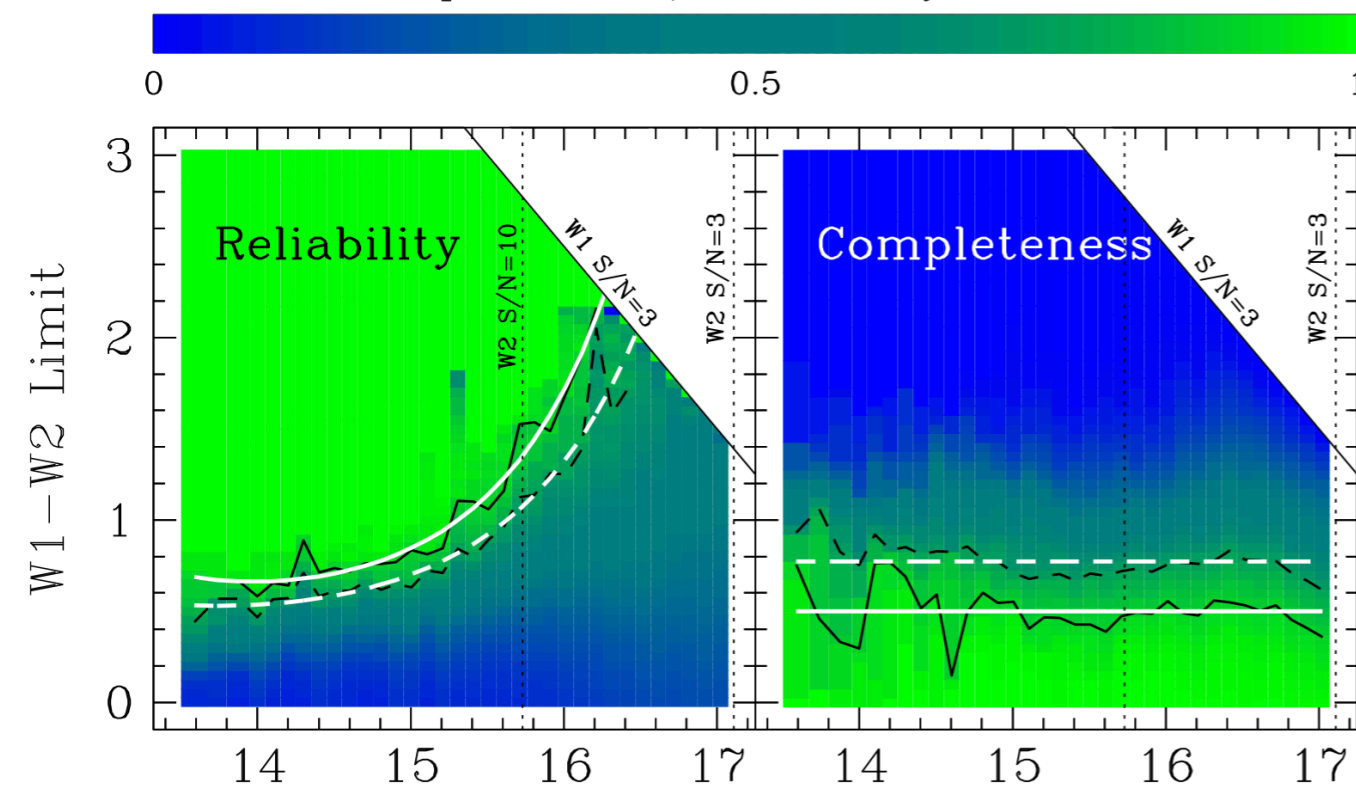
Target Selection



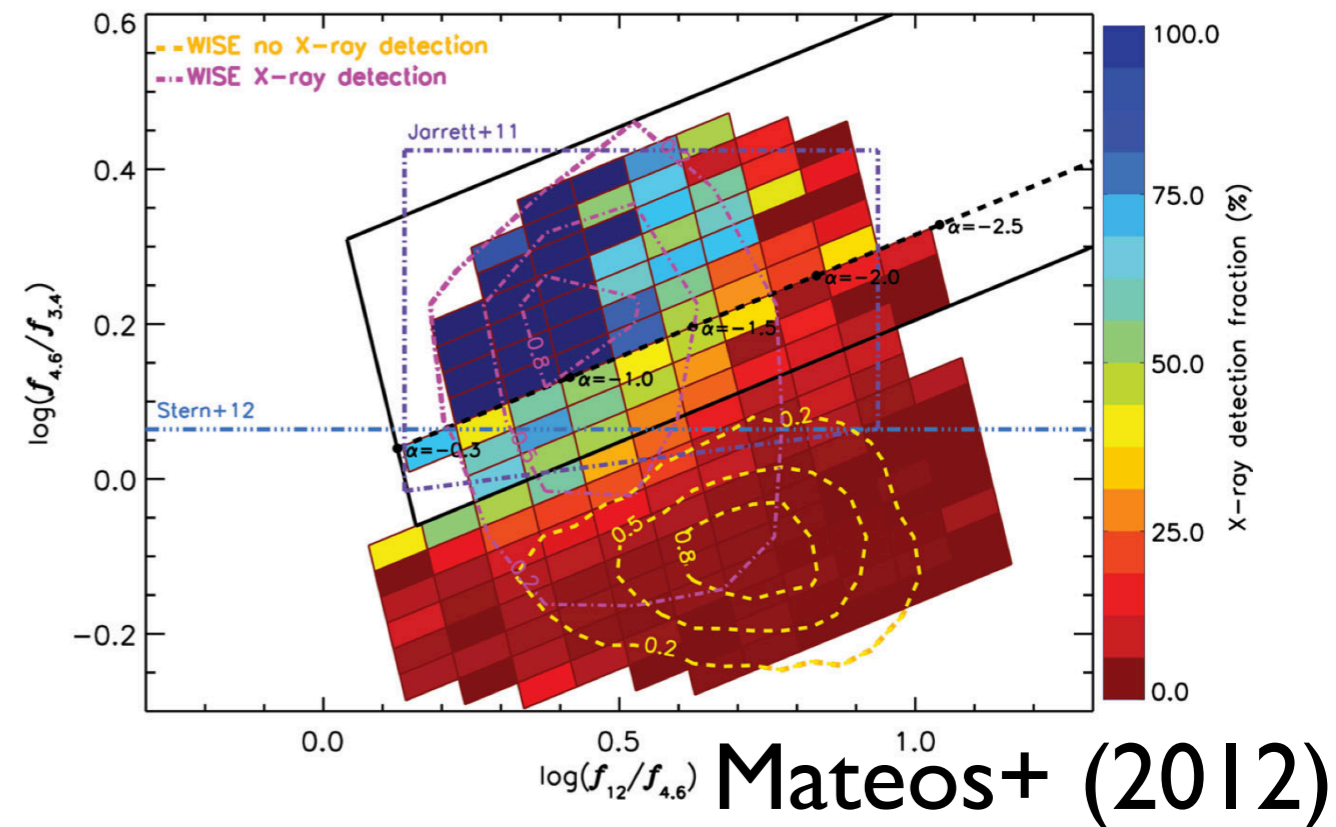
- Bayesian association method (Salvato, Buchner, et al.)
- An extension of Budavari & Szalay (2008) method
- Allows input priors based on any combination of counterpart characteristics
- But what are the best priors to use?
 - Must be available over full BOSS footprint
 - Must cope with heterogeneous counterpart types
- Many RASS sources already have SDSS spectrum (Anderson+, 2003, 2007)

- X-ray bright AGN are **red** in the MIR
- X-ray bright sources are **bright** in the MIR
- Field stars+galaxies are **faint** and/or **blue** in the MIR

Completeness/Reliability Fraction



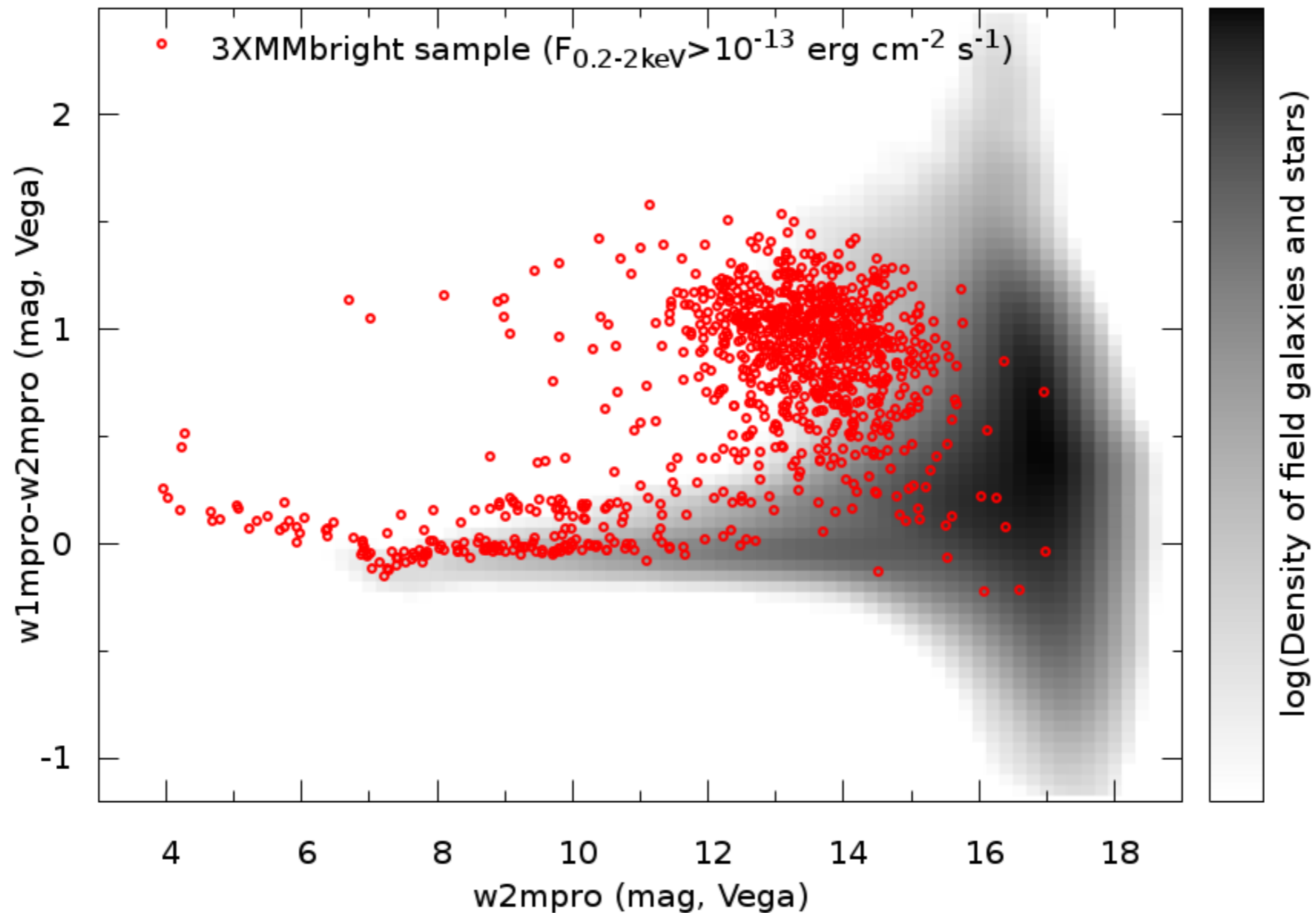
Assef+ (2013)



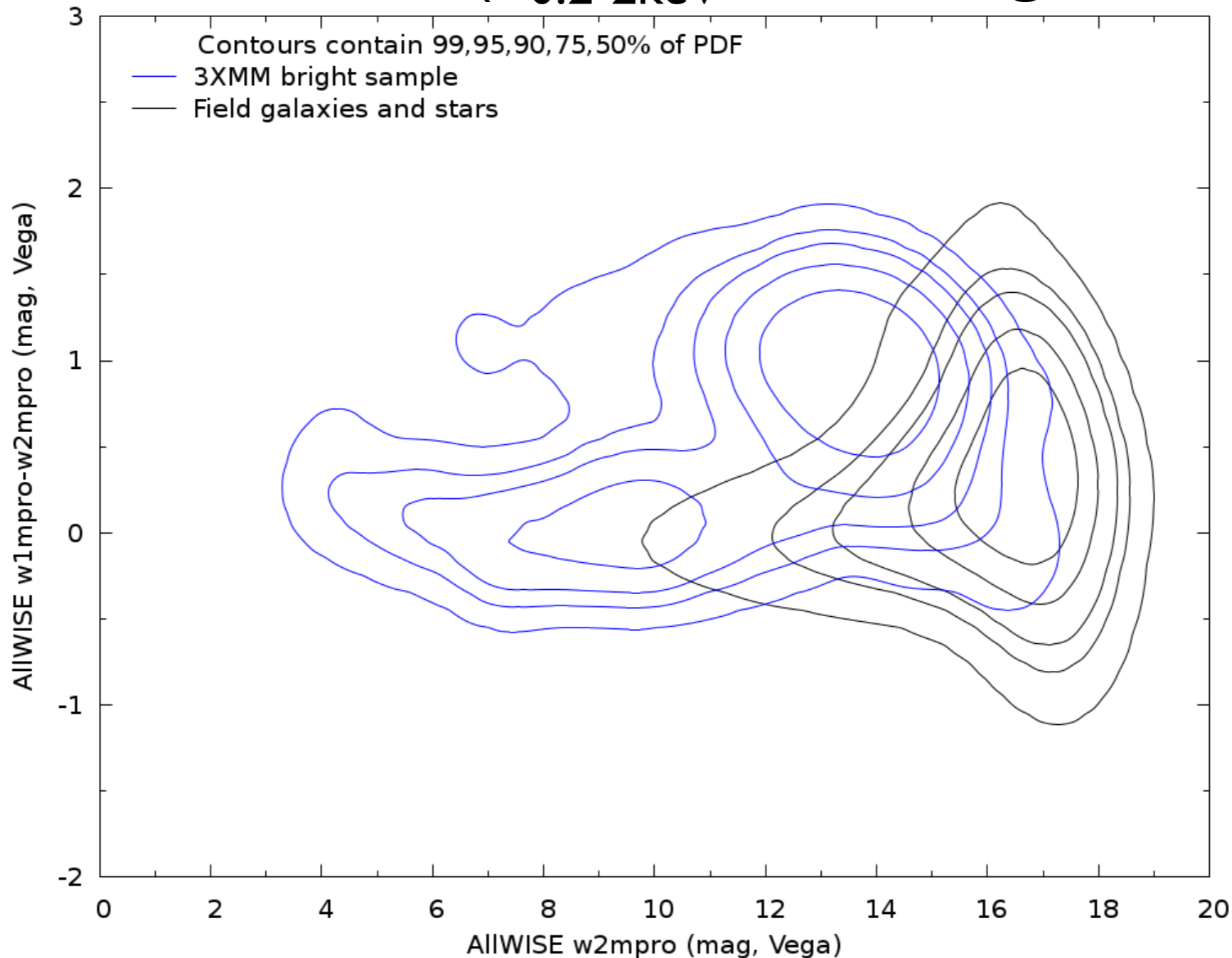
Mateos+ (2012)

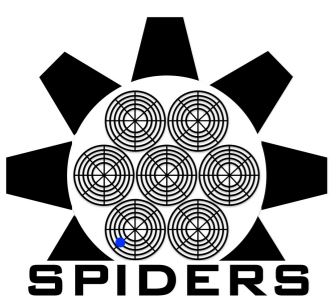
- Uses all of the WISE data from the primary mission
- 4 mid-IR bands ($\lambda_{\text{eff}} = 3.4, 4.6, 12$ and $22\mu\text{m}$)
- 95% of sky covered to 5σ limits of at least:
 - $3.4\mu\text{m}$ (W1) \rightarrow 17.6 Vega = 20.3 AB
 - $4.6\mu\text{m}$ (W2) \rightarrow 16.1 Vega = 19.5 AB
 - $12\mu\text{m}$ (W3) \rightarrow 11.5 Vega = 16.7 AB
- Unobscured QSOs typically have colours:
 $r' - W2 \sim 5$ (Vega, see e.g. DiPompeo+14)
- Obscured QSOs much redder

- Train on a sample of ~ 1000 bright point-like 3XMM sources ($F_{0.2-2\text{keV}} > 10^{-13} \text{ erg cm}^{-2} \text{ s}^{-1}$)



- Train on a sample of ~ 1000 bright point-like 3XMM sources ($F_{0.2-2\text{keV}} > 10^{-13} \text{ erg cm}^{-2} \text{ s}^{-1}$)

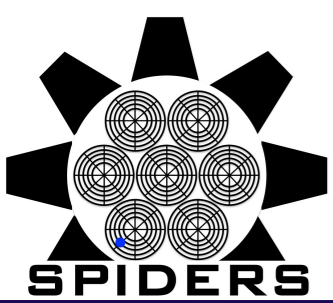




Target Selection



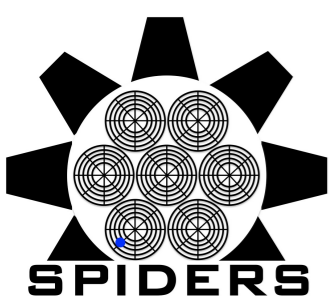
- 30855/32408 RASS sources have an AllWISE counterpart with $P \geq 0.01 \rightarrow$ choose maximum P
- 28515/30855 have ≥ 1 match in SDSS imaging ($< 1.5''$)
- Filter out all objects which:
 - already have SDSS spectra (11643 sources)
 - have $i'_{\text{fiber}} < 17$ or $i'_{\text{fiber}} > 22.5$ (7411+283 sources)
 - lie near very bright stars (150 sources)
- Leaves 9028 RASS AGN targets for eBOSS
- Similarly, 819 XMMSL AGN targets in eBOSS



X-match validation



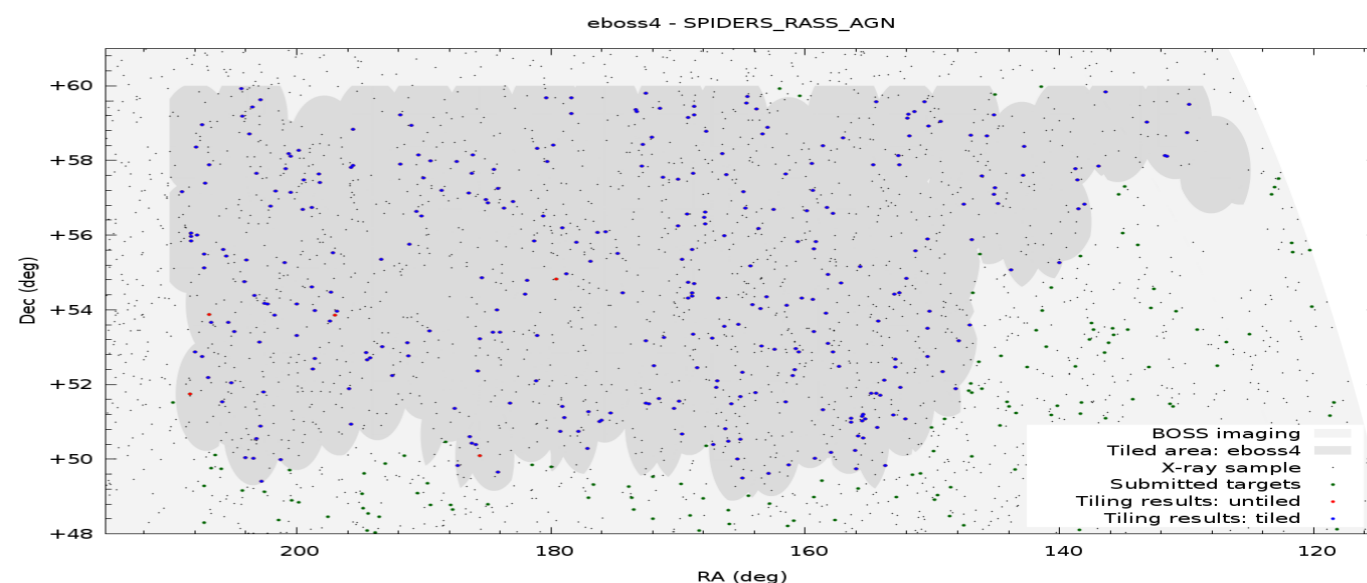
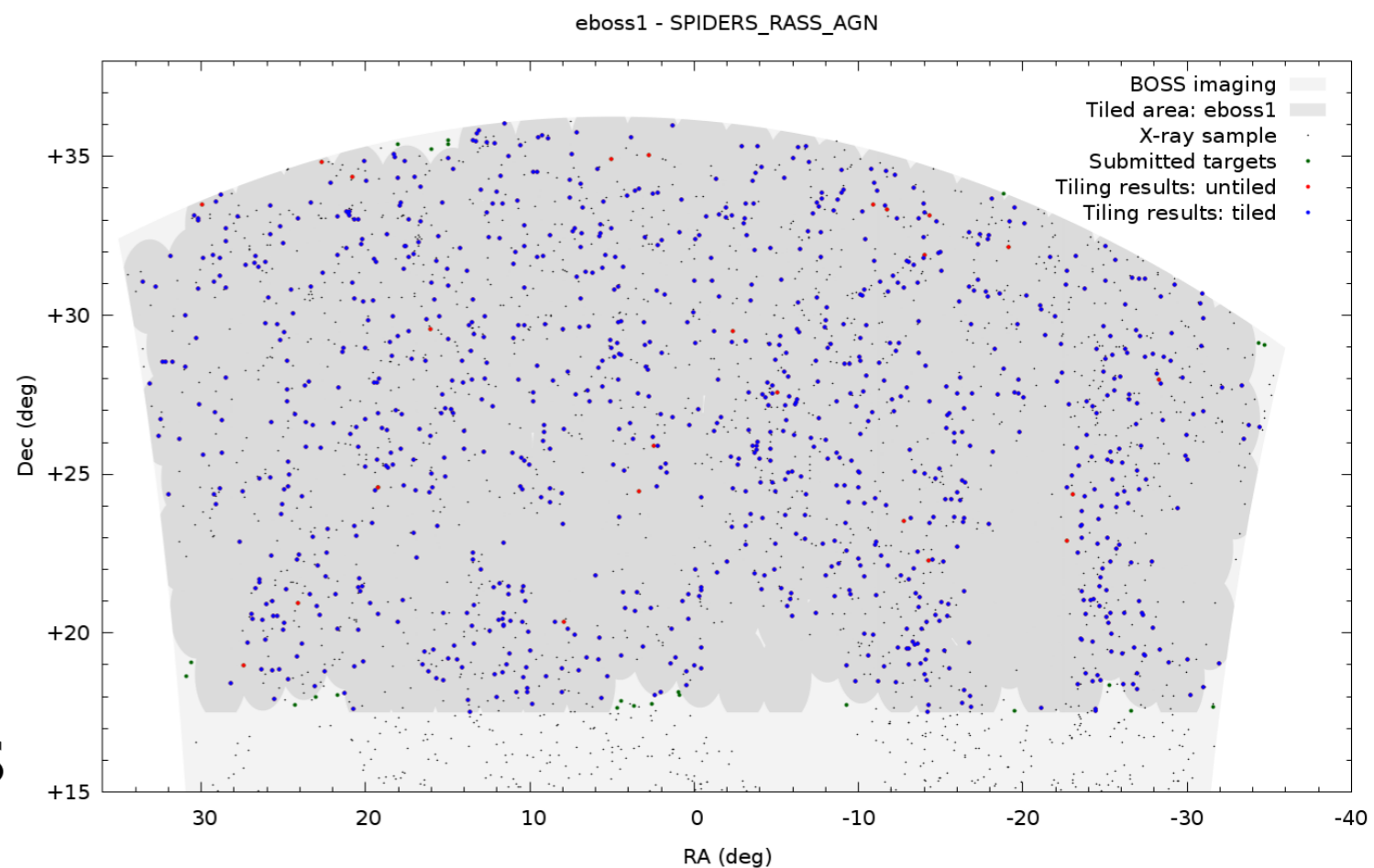
- How well do we recover correct counterparts when starting with RASS positions+errors?
- Use 3XMMBright→AllWISE matches as ‘truth’
 - 295/1000 of 3XMMBright+AllWISE sample are matched to RASS sources
 - For 279/295 we choose exactly the same WISE counterpart that was chosen using XMM position
 - = 94.6 ± 1.3 % success rate!
- Similar test against bright ISPXS (Swift-XRT) sources gives $940/1006 = 93.4 \pm 0.8$ % success rate

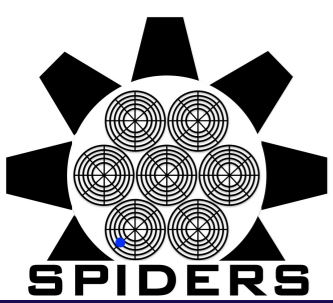


eBOSS tiling results



- Tiling results for first 4 chunks of eBOSS (2840deg²)
- SPIDERS targets tiled with high priority
- RASS AGN: 3084 targets
- XMMSL AGN:
 - 288 targets
- Over 98% of submitted targets will get spectra

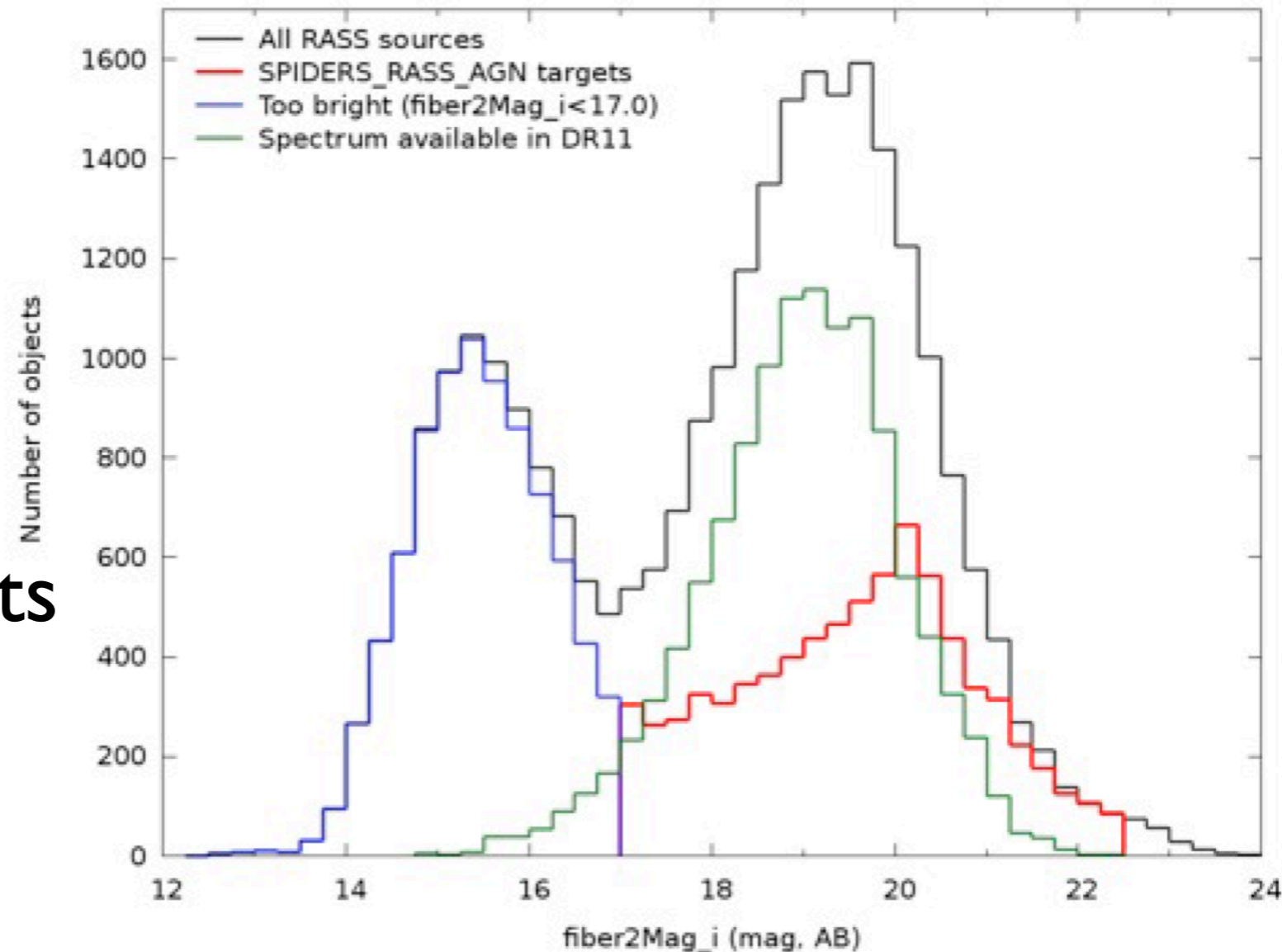




eBOSS tiling results

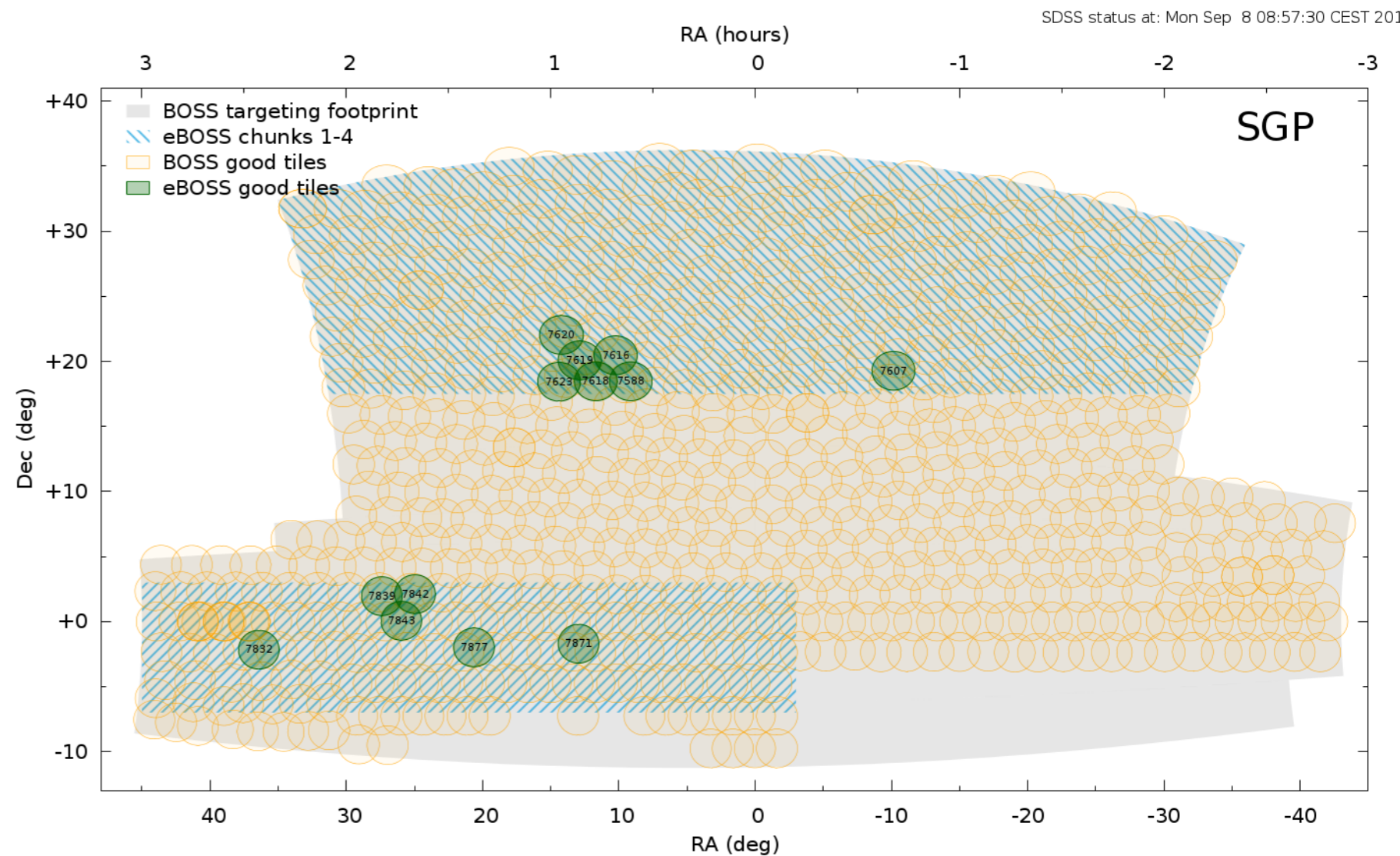


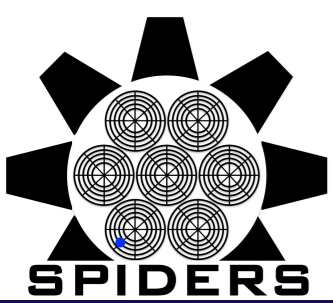
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- Over 98% of submitted targets will get spectra



- First 13 eBOSS plates have already been observed

PIPE-LINE CLASS	RASS AGN	XMMSL AGN
QSO	44	6
GALAXY	15	1
STAR	6	-
WARNING	1	-

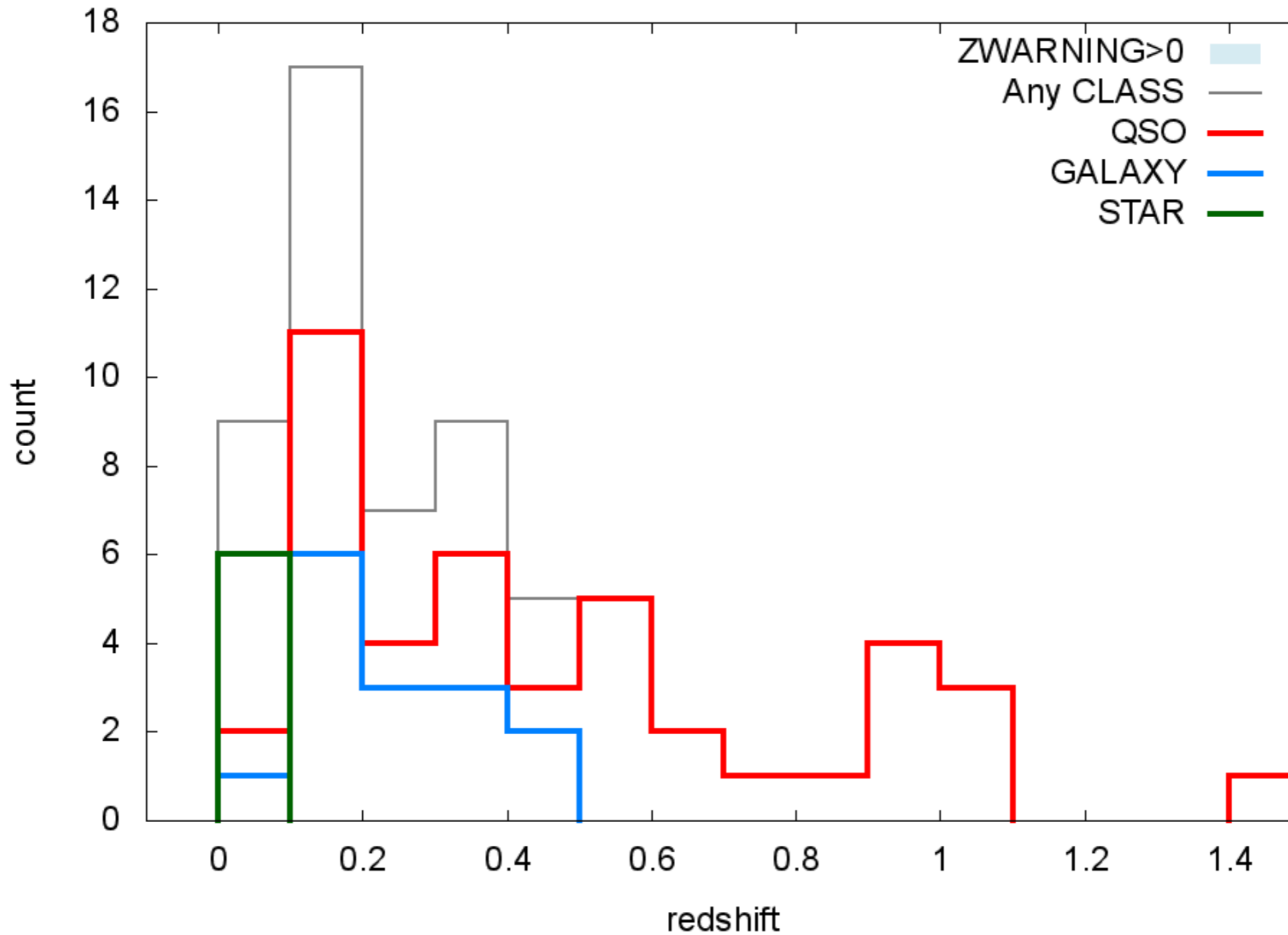


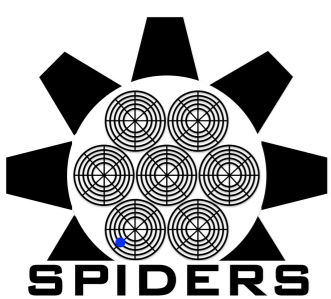


eBOSS first results



eBOSS data at: Mon Sep 8 08:57:30 CEST 2014



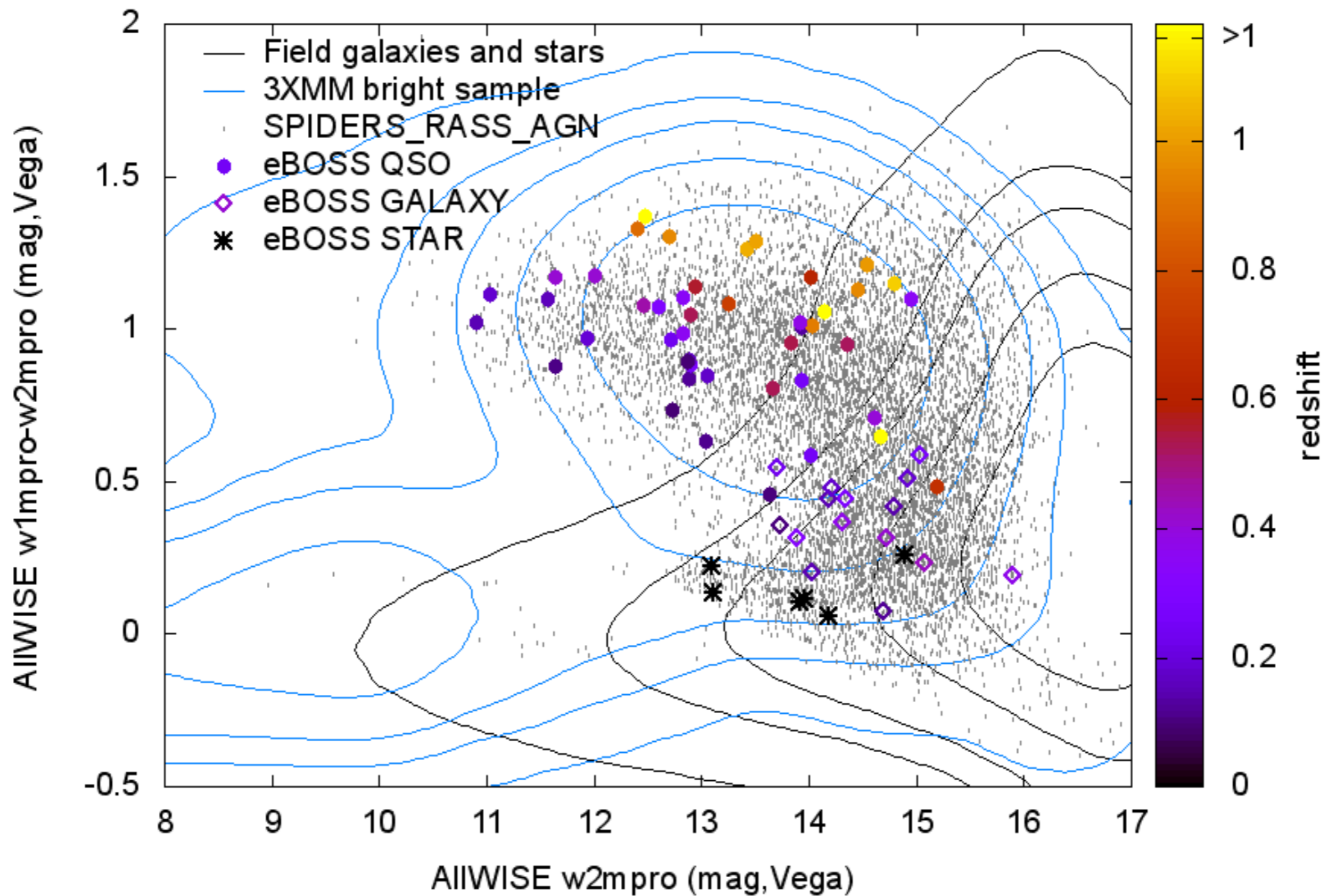


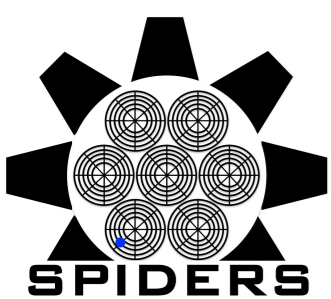
eBOSS first results



eBOSS data at: Mon Sep 8 08:57:30 CEST 2014

eBOSS redshifts for SPIDERS_RASS_AGN





eBOSS forecast



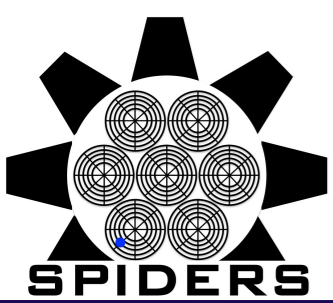
- Expected completeness for RASS AGN after combining eBOSS + SDSS-I+II+III
 - Fraction of RASS with ≥ 1 $P > 0.01$ WISE (0.952)
 - × Fraction with correct WISE counterpart (0.94)
 - × Fraction with SDSS counterpart (0.924★)
 - × Fraction with SDSS $i' < 22.5$ (0.983)
 - × Fraction outside bright star mask (0.991)
 - × Tiling completeness (0.981)
 - × Redshift success rate (~ 0.98)
- = 77% completeness, 94% reliability over 7500deg² inc. > 15000 X-ray selected AGN with redshifts

★ Lost fraction includes some saturated stars, and extended galaxies

Thanks

More details on the SDSS IV wiki:

[https://trac.sdss.org/wiki/eBOSS/Targets/SPIDERS/
AllWISETargetSelection](https://trac.sdss.org/wiki/eBOSS/Targets/SPIDERS/AllWISETargetSelection)



SEQUELS first results



- 66 SEQUELS plates observed
- RASS AGN targets selected using u'+r' priors (no WISE)

PIPE-LINE CLASS	RASS AGN
QSO	51
GALAXY	51
STAR	51
WARNING	10

