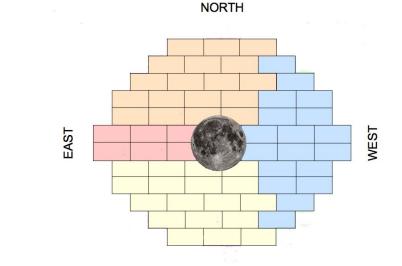


Dark Energy Camera

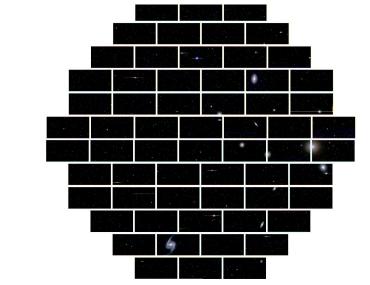
- 570 Mpixels
- 62 science CCDs
- 0".263 per pixel
- FOV: 2.2 deg, 3 deg²
- ugrizY filters (+VR, N964, and N662)
- Blanco 4m telescope

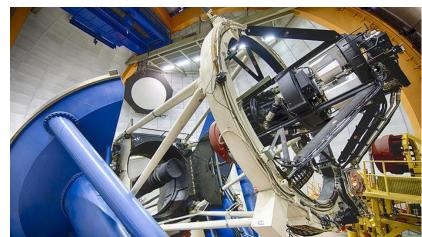




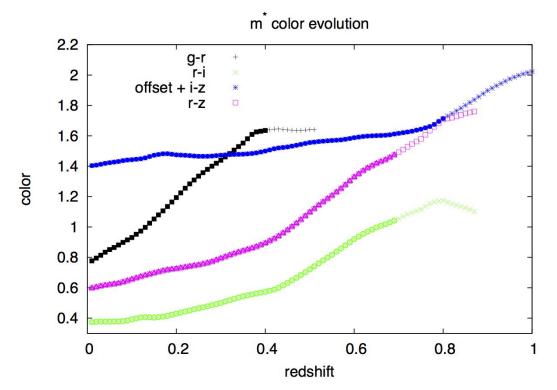
Dark Energy Camera

- Depth requirements for eROSITA photo-z
- MagLite I and II
- DECaLS + BlinK
- DES
- BLISS
- DECam eROSITA Survey





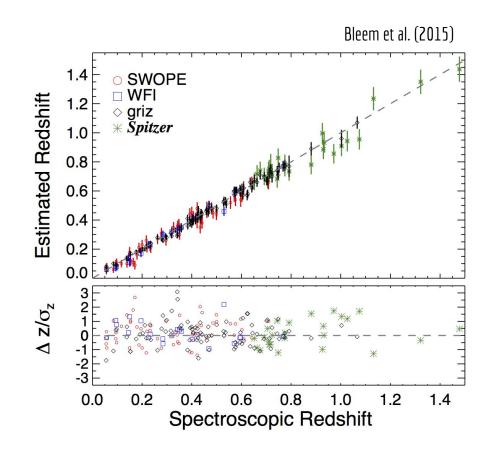
Depth requirements



- Red sequence photo-zs: griz
- Passively evolving stellar population model -> m*
- m*+1:
 - g=22.2, r=23, i=22.4, z=21.9 (to z=0.8 at 10 sigma)
 - tg=90s, tr=160, ti=110s, tz=140s
 - 100 deg² in griz under 8 hours.

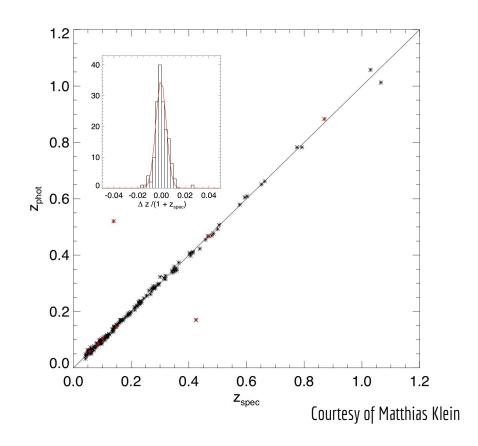
Depth requirements

- Red sequence photo-zs: griz
- Passively evolving stellar population model:
 - g=22.2, r=23, i=22.4,z=21.9
 (m*_{red}+1 to z=0.8 at 10 sigma)
 - tg=90s, tr=160, ti=110s, tz=140s
 - 100 deg² in griz under 8 hours.
- SPT followup



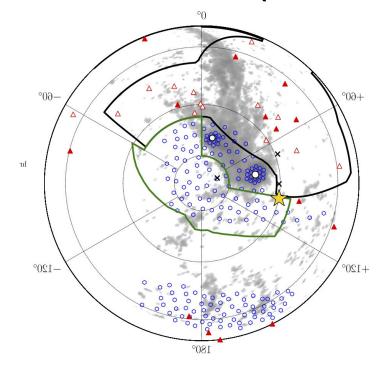
Depth requirements

- Red sequence photo-zs: griz
- Passively evolving stellar population model:
 - g=22.2, r=23, i=22.4,z=21.9
 (m*_{red}+1 to z=0.8 at 10 sigma)
 - tg=90s, tr=160, ti=110s, tz=140s
 - 100 deg² in griz under 8 hours.
- Matthias K. photoz will do even better



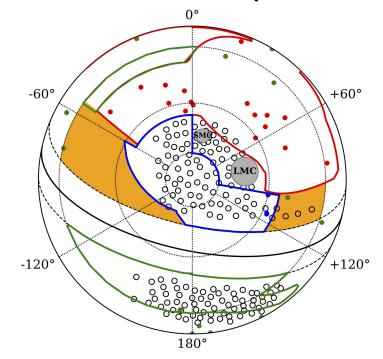
Surveys: "Magellanic Satellites Survey: The Search for Hierarchical Structures within the Local Group"

- MagLite I. PI: Keith Bechtol
- 12n, 2016A & 2017A
- 3x90s in *gr*
- 1300 sq-deg
- To search for ultra-faint galaxies and other low-surface-brightness stellar substructures associated with the Magellanic system
- Proprietary time: 1 Yr. (June 2018)
- Drlica-Wagner et al. (2016)



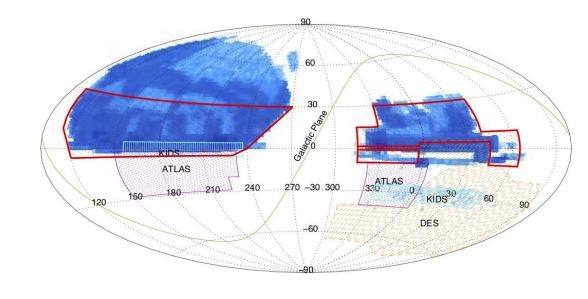
Surveys: "Magellanic Satellites Survey: The Search for Hierarchical Structures within the Local Group II"

- MagLite II. PI: Keith Bechtol
- 4n, July 2018A, ?n 2019A
- 3x90s in *gr*
- 1300 sq-deg
- To search for ultra-faint galaxies and other low-surface-brightness stellar substructures associated with the Magellanic system
- Proprietary time: 1.5 Yr. (Jan 2020)



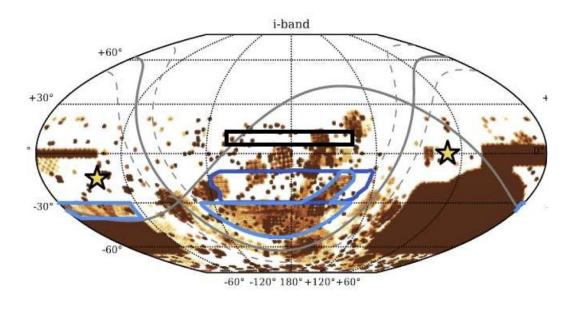
Surveys: "The DECam Legacy Survey of the SDSS Equatorial Sky (DECaLS)"

- PI: David Schlegel & Arjun Dey
- 157n (2014A-2018B)
- Decl. < 34 deg (NGC) & Decl. < 32 deg (SGC) . 9000 deg²
- 166s, 134s, and 200s in g, r, z.
- **g**=24, **r**=23.4, **z**=22.5 at 5sigma
- To provide targets for DESI
- Proprietary time: none



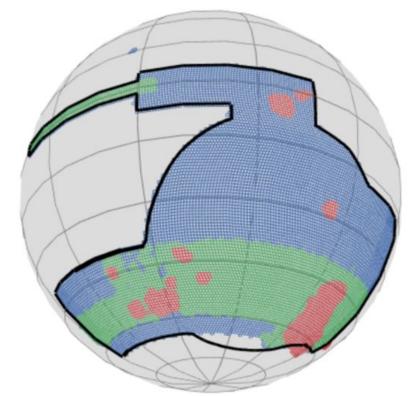
Surveys: "Blanco i-band Kilo-degree Survey (BlinK)"

- PI: Martin Makler
- 3n
- 600 deg²
- 3x90s *i-band*
- Follows the DECaLS footprint, partially
- Galaxy Clusters, Strong lensing, Transition galaxies, and Lyman break analogs.
- Proprietary time: 1.5 yrs. (Oct 2019)



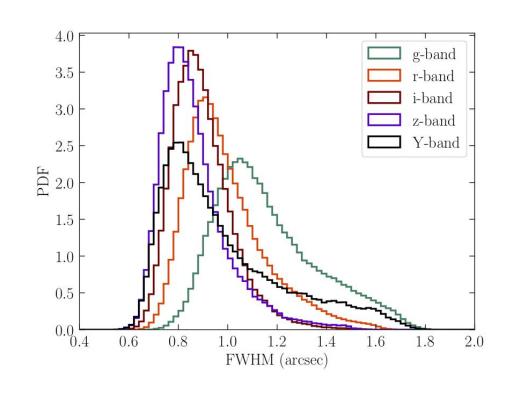
Surveys: "Dark Energy Survey (DES)"

- PI: Josh Frieman
- +525n (2012B-2018B)
- 5000 deg²
- grizY (90s in griz and 45s in Y)
- Cosmological probes
- Proprietary time: 1 Year



Surveys: "Dark Energy Survey (DES)"

- DR1: 345n
- Median PSF:
 - g=1.12", r=0.96", i=0.88", z=0.84", Y=0.90"
- Depth
 - g=24.33, r=24.08, i=23.44, z=22.69, Y=21.44
- 310M galaxies, 80M stars
- Abbott et al. (2018)
- https://des.ncsa.illinois.edu/releases/dr1

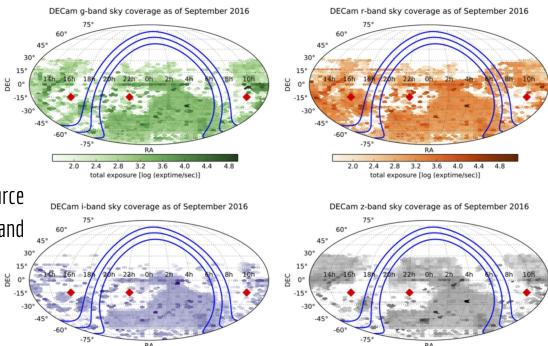


Surveys: "Blanco Imaging of the Southern Sky (BLISS)"

2.4 2.8 3.2 3.6 4.0 4.4

total exposure [log (exptime/sec)]

- PI: Marcelle Soares-Santos
- 11.5n
- -1000 deg^2 , decl < -30
- 3x90s *griz*
- Template images for Gravitational Waves source candidates, discovery of Milky way satellites, and search of Planet 9
- Proprietary time: None

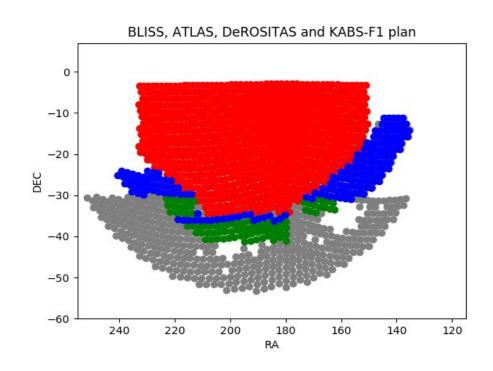


2.0 2.4 2.8 3.2 3.6 4.0 4.4 4.8 5.2

total exposure [log (exptime/sec)]

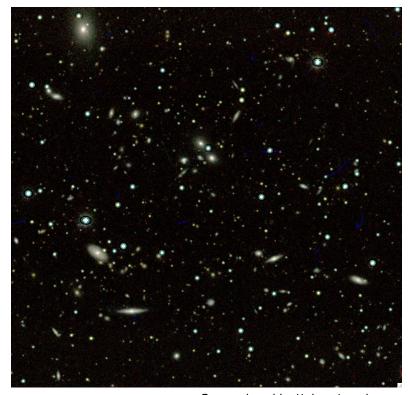
Surveys: "DECam eROSITA Survey (DeROSITAS)"

- 5 nights in 2017A:
 - 750 deg² in z band
 - 400 deg^2 in *gri* bands
- 5 nights in 2018A (starting in 5 days!)
- griz
 - g=22.2, r=23, i=22.4,z=21.9
 - Constraints: Fill the DE sky!
- As in 2017A we have engaged in collaborative efforts with other surveys to maximize the area coverage and the science return for the time invested

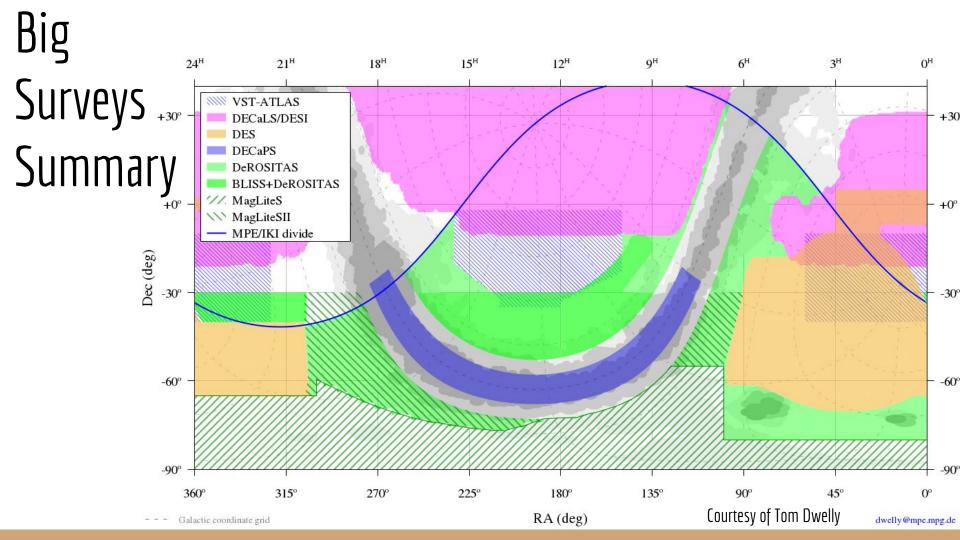


Surveys: "DECam eROSITA Survey (DeROSITAS)"

- 5 nights in 2017A:
 - $750 \text{ deg}^2 \text{ in } z \text{ band}$
 - 400 deg^2 in *gri* bands
- 5 nights in 2018A (starting in 5 days!)
- griz
 - g=22.2, r=23, i=22.4,z=21.9
 - g=23.5, r=23.3, i=22.7, z=22.1 (obs)
- DECaLS tiling
- Data reduction & photo-zs at Munich Sternwarte (Mohr, Holger, Klein et al.)



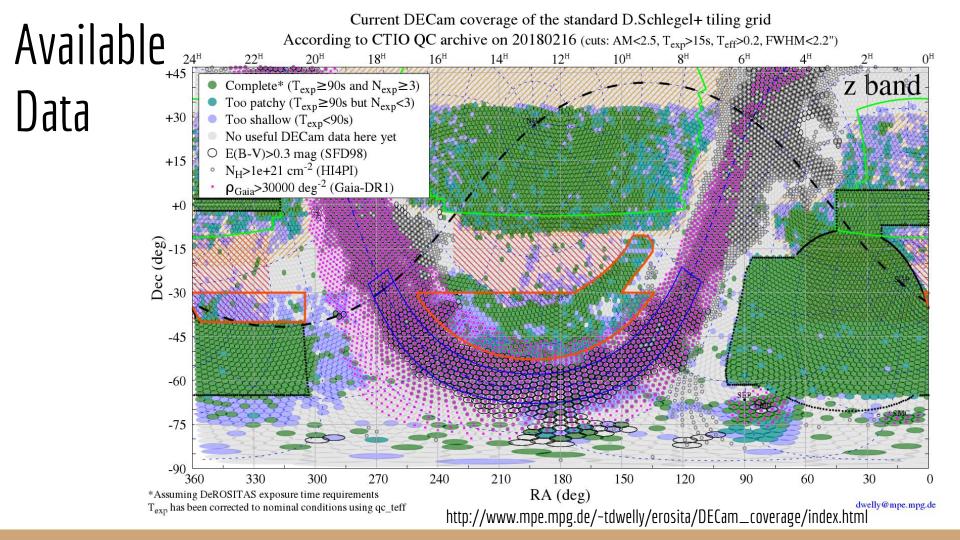
Data reduced by Holger Israel



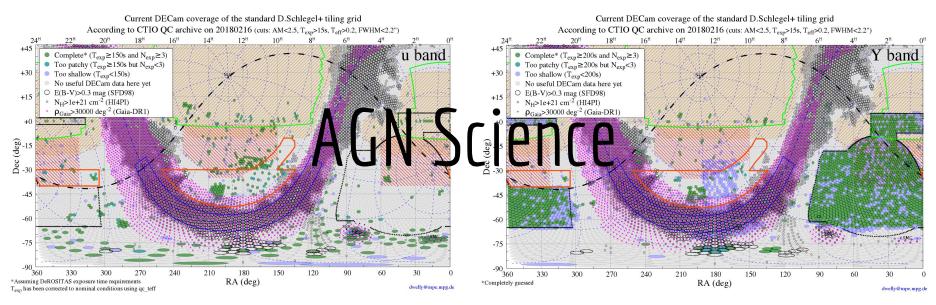
Current DECam coverage of the standard D.Schlegel+ tiling grid Available According to CTIO QC archive on 20180216 (cuts: AM<2.5, T_{exp}>15s, T_{eff}>0.2, FWHM<2.2") 14^{H} 10^{H} g band Complete* $(T_{exp} \ge 25s \text{ and } N_{exp} \ge 3)$ Data Too patchy ($T_{exp} \ge 25s$ but $N_{exp} < 3$) +30 Too shallow $(T_{exp} < 25s)$ No useful DECam data here yet E(B-V)>0.3 mag (SFD98) +15 $N_{H}>1e+21 \text{ cm}^{-2} \text{ (HI4PI)}$ ρ_{Gaia} >30000 deg⁻² (Gaia-DR1) +0 Dec (deg) -45 -60 -75 -90 <u>-</u> 360 330 300 270 240 210 180 150 120 90 60 30 RA (deg) *Assuming DeROSITAS exposure time requirements dwelly@mpe.mpg.de Texp has been corrected to nominal conditions using qc_teff http://www.mpe.mpg.de/-tdwelly/erosita/DECam_coverage/index.html

Current DECam coverage of the standard D.Schlegel+ tiling grid Available According to CTIO QC archive on 20180216 (cuts: AM<2.5, Texp>15s, Teff>0.2, FWHM<2.2") 16^{H} 14^{H} 10^{H} r band Complete* $(T_{exp} \ge 75s \text{ and } N_{exp} \ge 3)$ Data Too patchy ($T_{exp} \ge 75s$ but $N_{exp} < 3$) +30 Too shallow $(T_{exp} < 75s)$ No useful DECam data here yet E(B-V)>0.3 mag (SFD98) +15 $N_{H}>1e+21 \text{ cm}^{-2} \text{ (HI4PI)}$ ρ_{Gaia} >30000 deg⁻² (Gaia-DR1) +0 Oec (deg) 290 -30 -45 -60 -75 -90 330 300 270 240 210 180 150 120 90 60 30 RA (deg) *Assuming DeROSITAS exposure time requirements http://www.mpe.mpg.de/-tdwelly/erosita/DECam_coverage/index.html Texp has been corrected to nominal conditions using qc_teff

Current DECam coverage of the standard D.Schlegel+ tiling grid Available According to CTIO QC archive on 20180216 (cuts: AM<2.5, T_{exp}>15s, T_{eff}>0.2, FWHM<2.2") 14^{H} 10^{H} i band Complete* $(T_{exp} \ge 70s \text{ and } N_{exp} \ge 3)$ Data Too patchy ($T_{exp} \ge 70s$ but $N_{exp} < 3$) +30 Too shallow (T_{exp} <70s) No useful DECam data here yet \bigcirc E(B-V)>0.3 mag (SFD98) +15 $N_{H}>1e+21 \text{ cm}^{-2} \text{ (HI4PI)}$ $\rho_{\text{Gaia}} > 30000 \text{ deg}^{-2} \text{ (Gaia-DR1)}$ +0 Dec (deg) -45 -60 -75 -90 330 300 270 240 210 180 120 90 150 60 30 *Assuming DeROSITAS exposure time requirements RA (deg) dwelly@mpe.mpg.de Texp has been corrected to nominal conditions using qc_teff http://www.mpe.mpg.de/-tdwelly/erosita/DECam_coverage/index.html



Surveys: DECam u/Y band...



Barrientos, Shanks et al. (DES area), and Arevalo et al. (NGC; 2 out of the 5n in 2018A for Y) on 2019 on. More info in Tom's page http://www.mpe.mpg.de/-tdwelly/erosita/DECam_coverage/index.html

What is left to cover the eROSITA-DE Area?

Time required to complete 3-dithers g-band= 28 hours

```
# Time required to complete 3-dithers r-band= 36 hours
# Time required to complete 3-dithers i-band= 39 hours
# Time required to complete 3-dithers z-band= 48 hours
# Number of nights for the program with 3d griz = 15.1 days (at 10 hours/night)
# Time required to complete 3-dithers Y-band= 97 hours (t<sub>..</sub>=150s; u=22.6)
# Time required to complete 3-dithers u-band= 69 hours (t_v= 270s; Y=21.0)
# Number of nights for the program with 3d ugrizY = 31.7 days (at 10 hours/night)
```