

ART-XC telescope update

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eROSITA Consortium meeting, Garching 23 – 26/4, 2018

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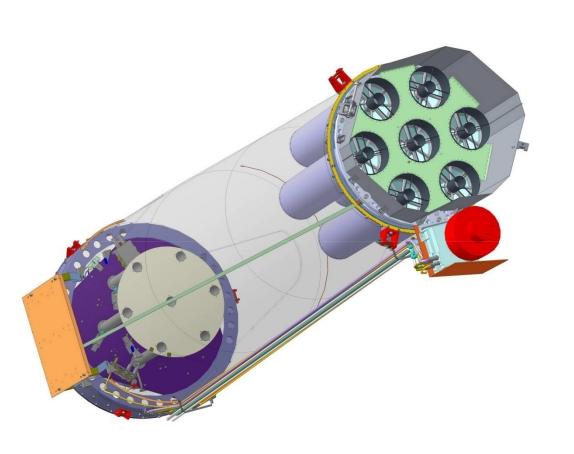
Last SRG news:

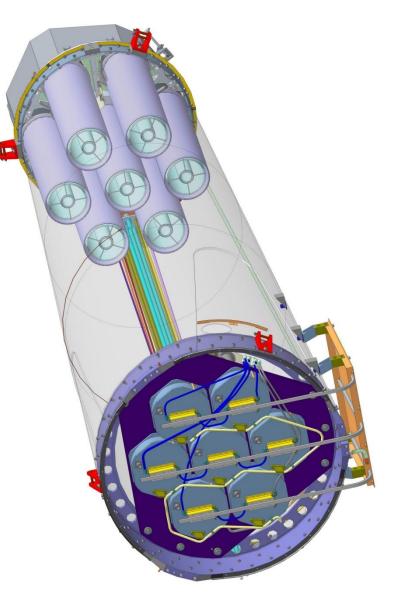
April 3, 2018 FM radiocomplex was delivered to Lavochkin Association from RKS

April 19, 2018 radiocomplex successfully passed entrance control at LA



ART-XC layout









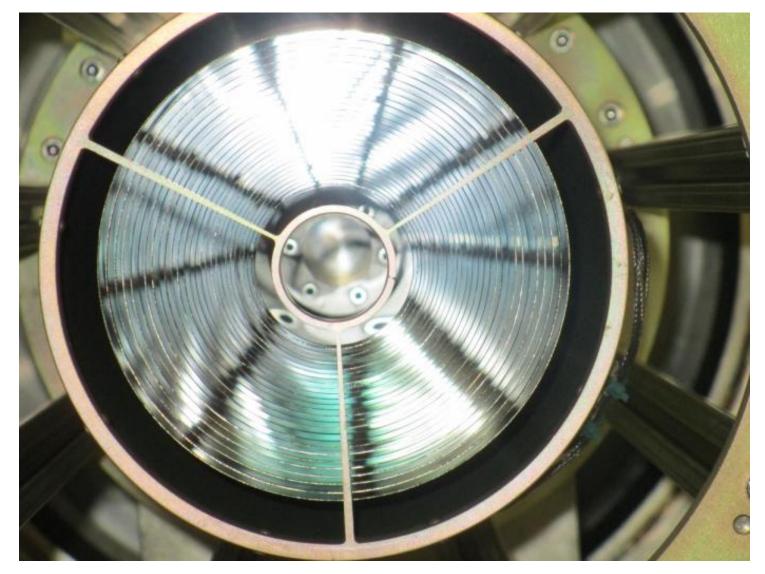
ART-XC: X-ray mirror systems MSFC/NASA

1.	Number of mirror systems	7
2.	Number of nested mirror shells	28
3.	Form of shell	Wolter-I
4.	FOV with detector Ø28.56 mm	~0.3 deg ²
5.	On-axis angular resolution, HPD	≤35″
6.	Focal length	2700 mm
7.	Length of shell	580 mm
8.	Diameter of mirror shells	49 – 145 mm
9.	Material of shells	Ni/Co
10.	Mirror coating materials	Iridium





ART-XC FM mirror system backside view

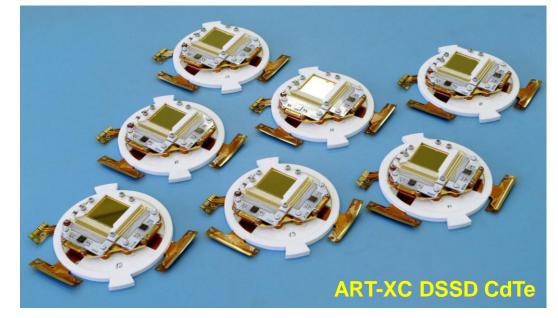




ART-XC: DSSD CdTe detector IKI RAS

- 1. CdTe manufacturerACRORAD
- 2. Dimensions $30 \times 30 \times 1 \text{ mm}^3$
- 3. Working area $28.56 \times 28.56 \text{ mm}^2$
- 4. Energy range 4 160 keV
- 5. Number of strips 48×48
- 6. Strip pitch 0.595 mm (45")
- 7. Be entrance window \emptyset 30 mm \times 100 μ m
- 8. ASIC, 2 pcs. VA64TA1
- 9. Dead time 0.77 ms
- **10. Working temperature 21° C**
- 11. Energy resolution $\leq 8.5\% @ 14 \text{ keV}$







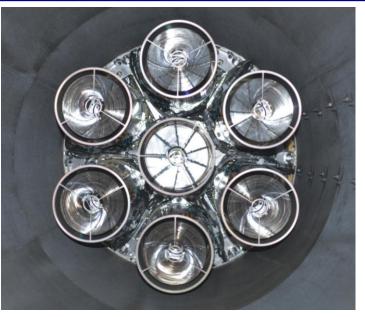






























- September 2017 April 2018 4 runs of ART-XC and S/C control system (BKU) software tests
- January 29 30, 2018 test of compatibilities of FM ART-XC and S/C reaction wheels \Rightarrow upper limit is 1000 rotations per minute (strong requirement)



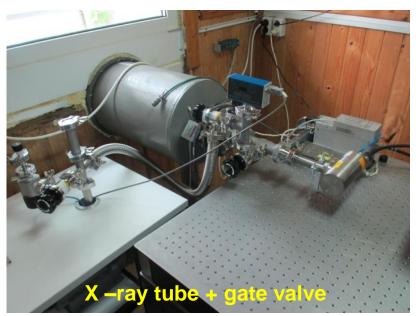
Calibration at IKI test facility mirror system + DSSD CdTe telescope ART-XC / SRG



IKI test facility: pipe 60 m + chamber 4 m X – ray source – mirror system distance: 61 m Residual pressure in pipe and chamber: $\leq 4 \times 10^{-5}$ bar X – ray tubes: Cr, Cu, Mo, Ag, 9 – 50 kV Al filters, 28 – 448 µm Two reference detectors – XR-100T-CdTe and XR-100SDD (Amptek) Two hexapods

<complex-block>







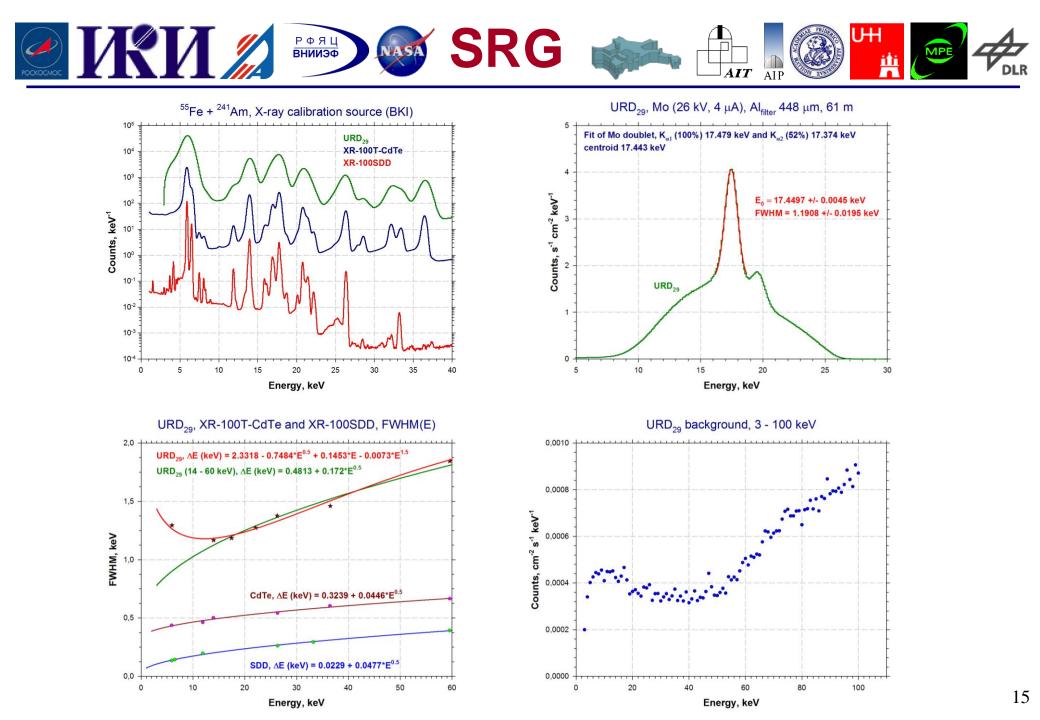






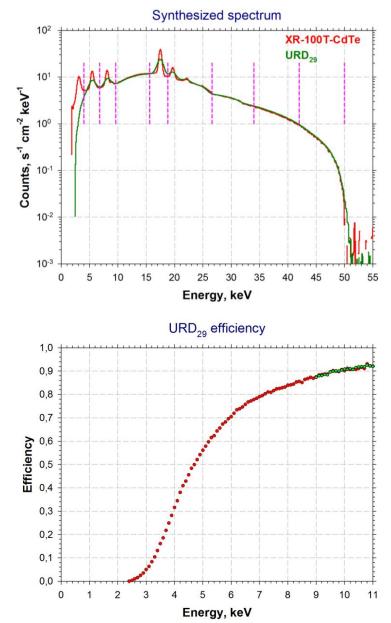




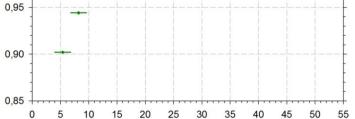




Ratio

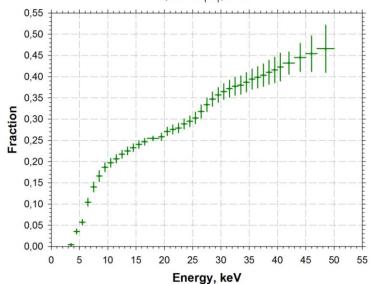


Ratio of fluxes, URD₂₉ / XR-100T-CdTe 1,10 1,05 1,00 0.95



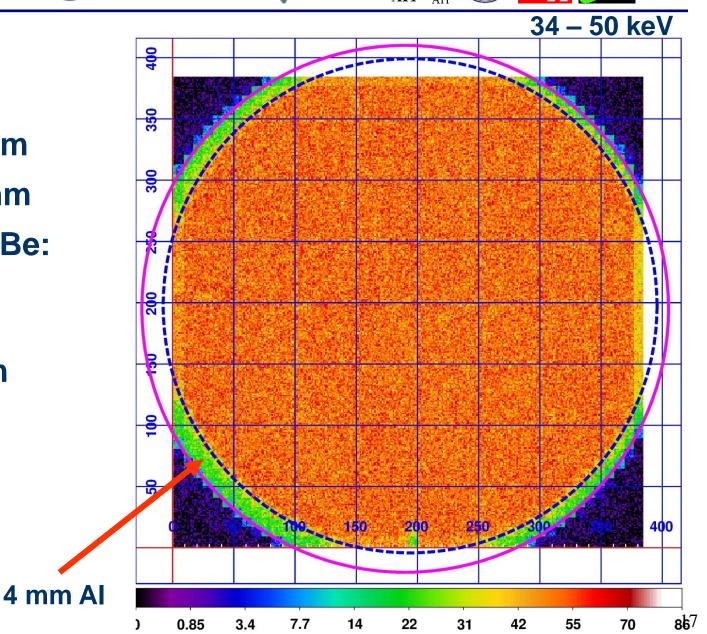
Energy, keV

 $(AII - B_1T_1) / AII$

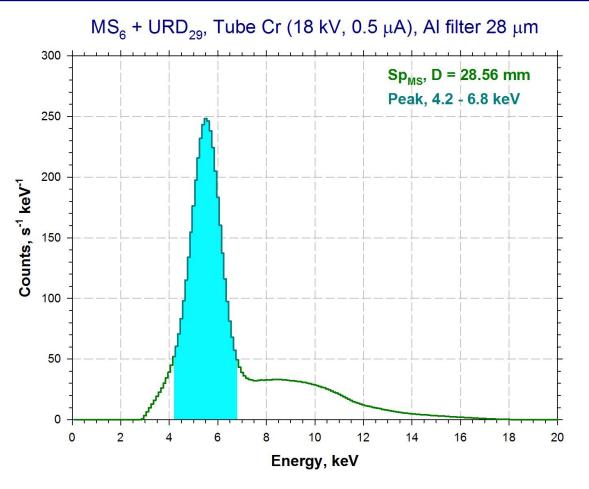




- URD₂₉, DSSD CdTe
- strips: **48**_X, **48**_Y
- strips pitch: 0.595 mm
- Be window: Ø30.0 mm
- Working area under Be: S = 6.896 cm²
- D_{int} of multilayer collimator : Ø32 mm



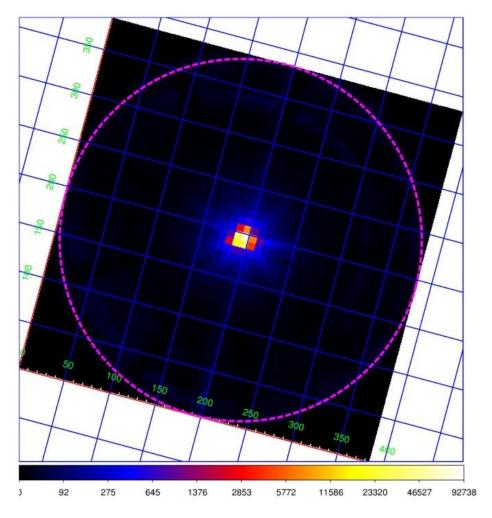


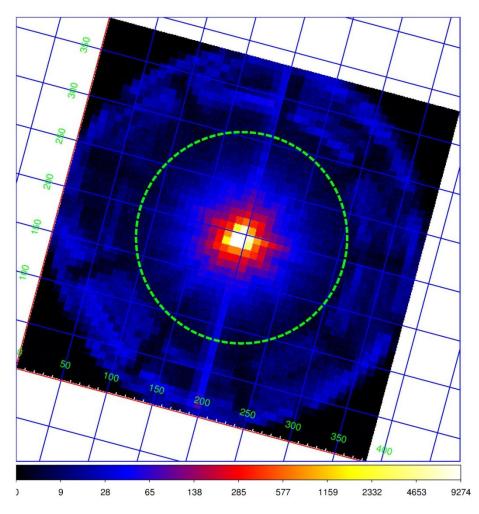


The integrated spectrum over an aperture \emptyset 28.56 mm, when the detector URD₂₉ is illuminated through the MS by an X-ray tube with a Cr anode (18 kV, 0.5 μ A, AI filter 28 μ m). The solid area shows the peak of the Cr lines in the range 4.2-6.8 keV.



MS₆ + URD₂₉, Cr (18 kV, 0.5 μA), AI 28 μm, 4.2 – 6.8 keV, 61 m



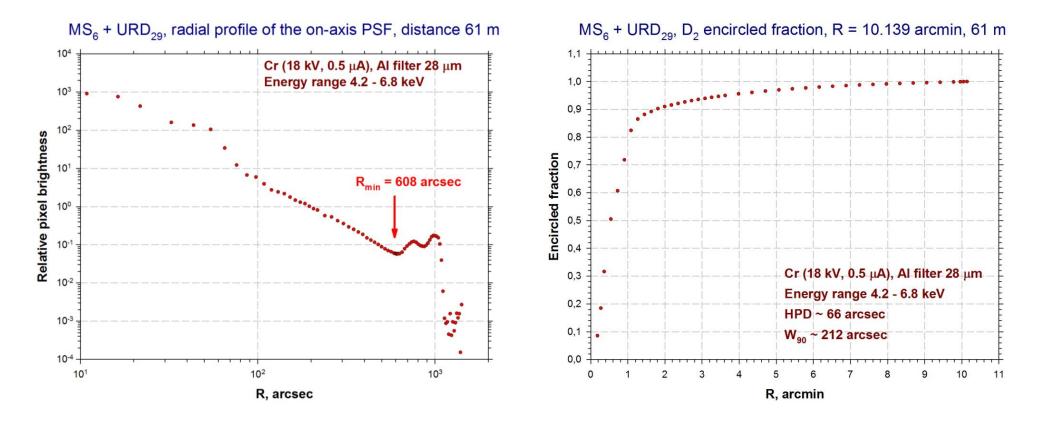


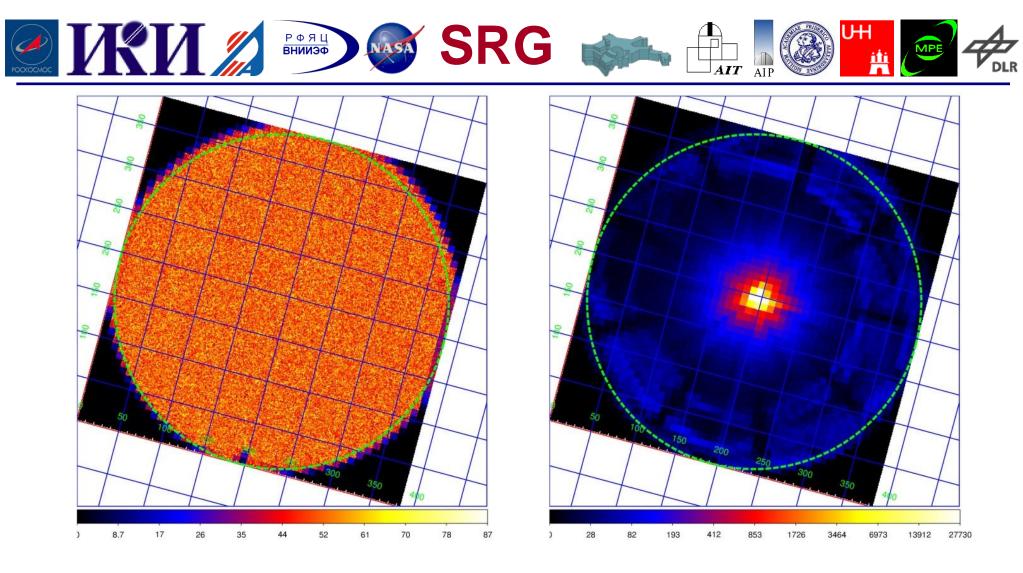
Logarithmic scale

10% of the maximum



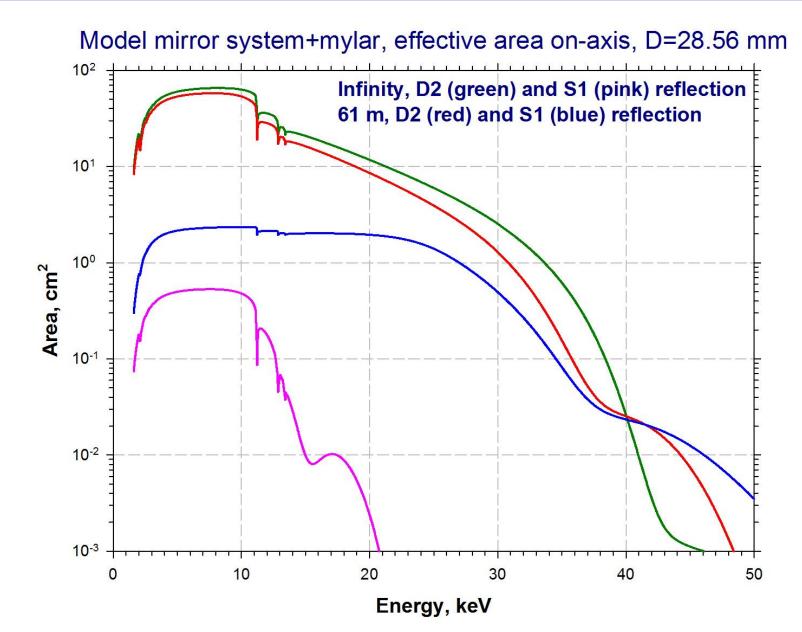
MS_6 + URD₂₉, Cr (18 kV, 0.5 μ A), Al 28 μ m, 4.2 – 6.8 keV, 61 m

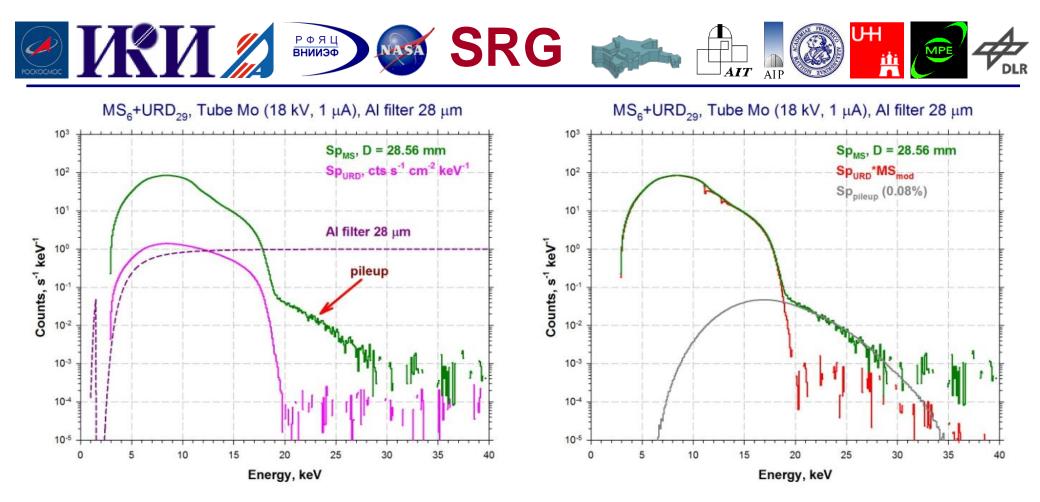




Left – detector image in the 4–18 keV band at direct illumination of URD₂₉ (linear color scale); Right – image of URD₂₉ illuminated through MS₆ (logarithmic color scale, cutoff at 10% of maximum value). The green dashed circle marks 28.56 mm diameter equal to ~34.76' for the 61 meter distance to the source.







Left: Sp_{URD} – integrated spectrum of events for direct (without MS) illumination of URD_{29} in units of cts s⁻¹ cm⁻² keV⁻¹ (pink); Sp_{MS} – integrated spectrum of URD_{29} illuminated through MS₆ within diameter of 28.56 mm (green); also shown here the photoabsorption transparency of the 28 µm thick AI filter (crimson). At energies above 19 keV in Sp_{MS} spectrum there is the pileup effect observed. Right: Sp_{MS} (green), $Sp_{URD} \times MS_{mod}$ – direct illumination spectrum multiplied by the mathematical model of the MS effective area MS_{mod} (red) and Sp_{pileup} – mathematical model spectrum describing the pileup effect, which relates to a small fraction ~0.08% of the total number of events (grey).



Thank you